

1 **Sedimentology and isotope geochemistry of transitional evaporitic**
2 **environments within arid continental settings: From erg to saline**
3 **lakes**

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10 **Short Title – Transitional evaporitic environments**

11 **ABSTRACT**

12 Arid continental basins typically contain a spectrum of coeval environments that coexist and interact
13 from proximal to distal. Within the distal portion, aeolian ergs often border playa, or perennial, desert
14 lakes, fed by fluvial incursions or elevated groundwaters. Evaporites are common features in these
15 dryland, siliciclastic dominant settings. However, sedimentary controls upon evaporite deposition are
16 not widely understood, especially within transitional zones between coeval clastic environments that
17 are dominantly controlled by larger scale allocyclic processes, such as climate. The sulphur ($\delta^{34}\text{S}$) and
18 oxygen ($\delta^{18}\text{O}$, $\Delta^{17}\text{O}$) isotope systematics of evaporites can reveal cryptic aspects of sedimentary cycling
19 and sulphate sources in dryland settings. However, due to the lack of sedimentological understanding
20 of evaporitic systems, isotopic data can be easily misinterpreted. This work presents detailed
21 sedimentological and petrographic observations, coupled with $\delta^{34}\text{S}$, $\delta^{18}\text{O}$ and $\Delta^{17}\text{O}$ data, for the early
22 Permian Cedar Mesa Sandstone Formation (western USA). Depositional models for mixed evaporitic /

23 clastic sedimentation, which occurs either in erg-marginal or lacustrine-marginal settings, are presented
24 to detail the sedimentary interactions present in terms of climate variations that control them.
25 Sedimentological and petrographical analysis of the evaporites within the Cedar Mesa Sandstone
26 Formation reveal a continental depositional environment and two end member depositional models have
27 been developed: erg-margin and lake-margin. The $\delta^{34}\text{S}$ values of gypsum deposits within the Cedar
28 Mesa Sandstone Formation are consistent with late Carboniferous to early Permian marine settings.
29 However, a marine interpretation is inconsistent with sedimentological and petrographic evidence.
30 Consequently, $\delta^{34}\text{S}$, $\delta^{18}\text{O}$ and $\Delta^{17}\text{O}$ values are probably recycled and do not reflect ocean-atmosphere
31 values at the time of evaporite precipitation. They are most likely derived from the weathering of older
32 marine evaporites in the hinterland. Thus, the results demonstrate the need for a combination of both
33 sedimentological and geochemical analysis of evaporitic systems to better understand their depositional
34 setting and conditions.

35 Keywords: Aeolian, Cedar Mesa Sandstone, gypsum, playa lakes, $\Delta^{17}\text{O}$, $\delta^{34}\text{S}$

36 **INTRODUCTION AND BACKGROUND**

37 Dryland continental basins typically contain a spectrum of depositional environments from alluvial fans
38 and immature fluvial systems within the proximal region (e.g. Blissenbach, 1954; Hooke, 1967; Bull,
39 1977; Blair & McPherson, 1994; Harvey *et al.*, 2005; Parsons & Abrahams, 2009), through to aeolian
40 and arid-fluvial systems within the medial (e.g. Langford, 1989; Langford & Chan, 1989; Herries, 1993;
41 Veiga *et al.*, 2002; Cain & Mountney, 2009; Al-Masrahy & Mountney, 2015; dos Reis *et al.*, 2019;
42 Formolo Ferronato *et al.*, 2019, Priddy & Clarke, 2020), and playa lake or lacustrine environments
43 within the distal region (Tunbridge, 1984; Huerta *et al.*, 2010). In the medial to distal regions of many
44 dryland continental basins, evaporite deposits are typically recognized along shallow lacustrine margins
45 (Carter & Pickerill, 1985; Orti *et al.*, 2003; Orti *et al.*, 2007), where they may form a distinct
46 depositional setting in their own right, or are dispersed within erg-marginal and ephemeral-fluvial
47 settings (Kocurek, 1988; Porter, 1986; Blakey, 2000; Clemmenson *et al.*, 2000; Tanner & Lucas, 2007).
48 However, evaporite deposits are not limited to only the distal regions within continental basins. Ancient

49 evaporite deposits have been described from proximal regions, typically deposited in association with
50 springs or alluvial fans in settings characterized by active tectonism and rapid subsidence (Southgate *et*
51 *al.*, 1989; Fischer & Roberts, 1991; Helvaci, 1995), and many modern examples, such as the acid saline
52 lakes of Western Australia, are deposited in topographic lows on cratons (e.g Aerts *et al.*, 2019).

53 The preserved clastic sediments of medial to distal aeolian and arid lake-marginal basinal settings are
54 recognized for their potential to act as high-quality hydrocarbon reservoirs or aquifers within the
55 subsurface (e.g. Glennie *et al.*, 1978), but interdigitated fluvial sediments are known to have detrimental
56 effects upon reservoir performance (e.g. North & Prosser, 1993). Consequently, the evolution of mixed
57 arid clastic systems, and the preservation of the sediments they deposit, have been well documented
58 (e.g. Langford & Chan, 1989; Herries, 1993; Veiga *et al.*, 2002; Mountney & Jagger, 2004). However,
59 the same cannot be said for mixed evaporitic–clastic systems of arid continental basins. Despite strong
60 associations between clastic and evaporitic environments, and between the deposits they produce,
61 evaporitic settings and their sediments are typically studied independently from the coeval siliciclastic
62 systems (Denison *et al.*, 1998;). Few studies examine the interdependence of one setting upon the other
63 and the complex sedimentary interactions between them. The studies that have addressed this issue
64 focus on the geochemistry and hydrology of the system (e.g. Huerta *et al.*, 2010) or look purely at the
65 sediments (e.g. Benison & Goldstein 2000; Gündogan *et al.*, 2005). Rarely do these studies take a
66 holistic interdisciplinary approach necessary for a complete understanding of these systems. This is
67 despite the obvious implications that inter-dispersed clastic sediments have upon the economic potential
68 of evaporitic deposits themselves, or the detrimental impacts that inter-dispersed evaporitic sediments
69 have upon clastic reservoirs and aquifers (Warren, 1999, 2006).

70 Evaporites can preserve geochemical proxies for palaeoenvironmental conditions, palaeoclimate and
71 evidence of past life that are typically absent from, or poorly preserved in, the siliciclastic deposits of
72 arid coeval settings. These include proxies for: palaeo-water and palaeo-air temperatures (Kovalevich,
73 1975; Petrichenko, 1979; Lowenstein & Spencer 1990; Goldstein & Reynolds, 1994; Roberts &
74 Spencer, 1995; Benison & Goldstein, 1999), ancient water chemistry (Horita *et al.*, 2002; Lowenstein

75 *et al.*, 2005) and atmospheric oxygenation (Blamey *et al.*, 2016; Blättler *et al.*, 2018, Crockford *et al.*,
76 2019).

77 The triple oxygen isotope ($\Delta^{17}\text{O}$) record, preserved in sedimentary sulphates, can be used to constrain
78 the coupled evolution of the biosphere and the atmosphere over a geological timescale (Crockford *et al.*
79 *et al.*, 2016; 2019) and may be the only geochemical proxy of atmospheric change (Bao, 2015).
80 Tropospheric oxygen (O_2) possesses a mass-independent, negative $\Delta^{17}\text{O}$ value, which is inherited from
81 ozone creation and destruction reactions in the stratosphere (Thiemens, 2006). The magnitude of the
82 negative $\Delta^{17}\text{O}$ value in tropospheric O_2 depends upon atmospheric composition (CO_2 and O_2
83 concentrations), as well as levels of gross primary productivity (Yung *et al.*, 1991; Crockford *et al.*,
84 2018). Detailed geochemical modelling can be used to deconvolve which factors principally control the
85 atmospheric $\Delta^{17}\text{O}$ value over Earth history (e.g. Waldeck *et al.*, 2019), but all approaches rest on the
86 assumption that the $\Delta^{17}\text{O}$ preserved in sedimentary sulphates is representative of the time period in
87 question.

88 The $\Delta^{17}\text{O}$ value of atmospheric oxygen is transferred to the sulphate anion complex during oxidative
89 sulphide weathering (Killingsworth *et al.*, 2018). It has been estimated that *ca* 21 to 34% of the oxygen
90 atoms in sulphate are derived from atmospheric O_2 (Kohl & Bao, 2011), although other estimates place
91 this as high as 60% (see Waldeck *et al.*, 2019, and references therein) and as low as *ca* 7% (Thurston *et al.*
92 *et al.*, 2010). Once formed, the sulphate complex does not readily undergo oxygen isotope exchange with
93 other fluids, other than under highly acidic, or high pressure/temperature, conditions (Kusakabe &
94 Robinson, 1977; Chiba *et al.*, 1981). As such, it is non-labile and considered to be stable through
95 relatively shallow, low-temperature diagenesis (Bao, 2015).

96 Oxygen isotope exchange with water, however, does occur as a result of microbial sulphur cycling,
97 which is prevalent in the marine (Turchyn & Shrag, 2004) and, as is being increasingly recognized, in
98 the continental realm (Benison & Bowen, 2013; Johnson *et al.*, 2015). Although previous studies have
99 focused mainly on tracking sulphur through microbial sulphur cycling, the fate of sulphate-bound
100 oxygen – particularly any mass-independent $\Delta^{17}\text{O}$ signal it may carry – during sulphide oxidation and

101 microbial sulphate cycling has been a more recent development (Hemingway *et al.*, 2020). Although
102 sulphates precipitated from marine waters do carry a significant negative $\Delta^{17}\text{O}$ value, due to the
103 propensity of microbial cycling in the marine realm, it is generally acknowledged that the $\Delta^{17}\text{O}$ value
104 preserved by marine sulphates is a conservative estimate of the true, atmospheric $\Delta^{17}\text{O}$ value of the time
105 period (Crockford *et al.*, 2019).

106 The fluvial sulphate flux constitutes the sole input into marine sulphate reservoirs (Halevy *et al.*, 2012),
107 is derived from the oxidative weathering of pyrite (incorporating a contemporaneous atmospheric $\Delta^{17}\text{O}$
108 signal) and dissolution of older calcium sulphates (Halevy *et al.*, 2012; Wortmann & Paytan, 2012).
109 Thus, in order to use $\Delta^{17}\text{O}$ values to infer temporal changes in coupled atmosphere–biosphere evolution,
110 it is necessary to minimise the spatial variability in $\Delta^{17}\text{O}$ preservation as far as possible. As such,
111 continental, dryland evaporite sequences could prove an important inventory of atmospherically derived
112 $\Delta^{17}\text{O}$.

113 This study investigates the environmental, spatial and temporal relationships between evaporitic and
114 clastic sediments within the arid and dominantly clastic continental settings preserved within the Cedar
115 Mesa Sandstone Formation, south-eastern Utah, USA. It describes and interprets multi-scale
116 relationships between evaporitic and clastic sediments using an outcrop lithofacies analysis augmented
117 by petrographic and XRD data. Triple oxygen and sulphur isotope analyses, conducted on evaporite
118 samples, provide further insight into possible depositional controls upon the system and demonstrate
119 the value of examining evaporitic deposits in the context of their coeval clastic setting. Generic
120 depositional models that account for the observations within a clastic–evaporitic depositional setting
121 are developed, from which the possible larger scale allocyclic controls and their effects upon the
122 preserved deposits are discussed.

123 **GEOLOGICAL SETTING AND PREVIOUS WORK**

124 The early Permian (Cisuralian) Cedar Mesa Sandstone Formation of the Western USA was deposited
125 within the Paradox Basin: an oval shaped, Carboniferous flexural foreland basin formed by loading in

126 response to the uplift of the Ancestral Rocky Mountains (ARM), and defined in shape by the
127 depositional extent of the evaporites of the Paradox Formation (Mallory, 1960; Condon, 1997; Barbeau,
128 2003).

129 Exposed across much of southern Utah (Fig. 1), the dominantly clastic sediments of the Cedar Mesa
130 Sandstone Formation form part of a 4 km thick late Pennsylvanian to mid-Permian basin fill that was
131 derived principally from the Uncompahgre Uplift of the Ancestral Rocky Mountains that were present
132 to the north-east of the basin (Loope, 1984). The age of the formation is given by U-Pb detrital zircon
133 studies (Dickinson & Gehrels, 2003), vertebrate fossils and trackways (Lockley & Madsen Jr, 2008;
134 Gay *et al.*, 2020) and lateral and vertical relationships with other formations and coeval deposits that
135 can be dated from well constrained palaeontological data (Condon, 1997; Lucas & Krainer, 2005).
136 Originally described as the deposits of an aeolian system (McKnight, 1940; Baker, 1946), the formation
137 was later reinterpreted to be of shallow marine origin from the presence of broken shell fragments in
138 the cross-stratification (Baars, 1962; 1979; Mack, 1977; 1979). However, from grainfall–grainflow
139 couplets that form foresets of large-scale cross-stratification (Loope, 1984), and from further continental
140 indicators including rhizoliths (Stanescio & Campbell, 1989), an aeolian origin for the sediments was
141 reaffirmed by the 1980s (Loope, 1984) and is widely accepted today (Lanford & Chan, 1989; Mountney
142 & Jagger, 2004).

143 The Cedar Mesa Sandstone Formation is one of four lithostratigraphical units that comprise the Cutler
144 Group in the distal portion of the Paradox Basin. The oldest unit of the Cutler Group – the lower Cutler
145 beds – is an informal lithostratigraphical subdivision that comprises aeolian, fluvial and shallow marine
146 sediments (Jordan & Mountney, 2010) related to repeated cyclic transgressions driven by alternations
147 between an arid and humid climate (Jordan & Mountney, 2012). The lower Cutler beds are overlain
148 conformably by the predominantly wet aeolian sediments of the Cedar Mesa Sandstone Formation
149 (Mountney & Jagger, 2004), that are, in turn, conformably overlain by the terminal fluvial fan sediments
150 of the Organ Rock Formation (Cain & Mountney, 2009). Finally, aeolian sediments of the De Chelly
151 Sandstone form the youngest division of the Cutler Group in the distal parts of the basin (Blakey, 1996;
152 Dubiel *et al.*, 1996; Condon, 1997; Stanescio *et al.*, 2000).

153 To the north-east of the basin (Fig. 1A), sediments of the Cedar Mesa Sandstone Formation have a
154 complex interfingering relationship with coeval sediments of the Cutler Group that were shed south-
155 westerly into the basin from the Uncompaghre Uplift to form proximal to medial alluvial fan and fluvial
156 deposits (Mack, 1977; Mountney & Jagger, 2004). To the north-west of the basin, aeolian sediments of
157 the Cedar Mesa Sandstone interfinger with shallow-marine limestone indicating the location of the
158 palaeoshoreline (Fig. 1) (Loope, 1984).

159 The aeolian sediments of the Cedar Mesa Sandstone represent the deposits of a north-east – south-west
160 trending coastal erg system that developed along the shoreline, and was supplied with sediment sourced
161 from the local marine shelf (Loope, 1984; Blakey, 1988; Blakey *et al.*, 1988). The erg extended 100 km
162 south, with sediments of the erg centre preserved 130 km west of Blanding near Hite (Fig.1), in south-
163 eastern Utah (Langford & Chan, 1993; Mountney, 2006). Major flooding surfaces that are preserved in
164 the aeolian strata to the south-east of Canyonlands National Park (Fig.1) subdivide the erg sediments as
165 the dune-fields became progressively smaller and more isolated away from the erg centre (Langford &
166 Chan, 1989). Major dune field sediments of the erg are separated by the sediments of wet interdunes,
167 which grade into sabkha-like evaporitic deposits towards the south-east (Fig.1) (Blakey, 1988; Blakey
168 *et al.*, 1988; Peterson, 1988; Condon, 1997; Huntoon *et al.*, 2000; Mountney & Jagger, 2004; Langford
169 & Massad, 2014; Pettigrew *et al.*, 2019).

170 Sabkha-like evaporitic deposits of the Cedar Mesa Sandstone Formation are exposed around the town
171 of Bluff in south-eastern Utah. They show sedimentary features including nodular evaporites and
172 enterolithic growth structures interbedded with aeolian sands and fine-grained clastic sediments. These
173 strata have been interpreted as the product of deposition in inland sabkhas associated with playa lakes
174 of the coeval aeolian erg (e.g. Glennie, 1972), but sulphur isotope values ($\delta^{34}\text{S}$) of gypsum from the
175 succession fall within the narrow marine range of Permian $\delta^{34}\text{S}$ values (Claypool *et al.*, 1980; Stanesco
176 & Campbell, 1989) and suggest a marine influence on the depositional environment. However, carbon
177 ($\delta^{13}\text{C}$) and oxygen ($\delta^{18}\text{O}$) analysis of limestones within the succession only partly fall within the
178 common marine range (Hudson, 1977). Consequently, the evaporitic sediments of the Cedar Mesa
179 Sandstone were interpreted to be the product of a sabkha environment that was, in part, fed by marine

180 waters via unspecified methods, with mixing of fresh and marine waters under conditions of intense
181 evaporation (Stanescio & Campbell, 1989).

182 Recent work suggests an environment isolated from marine waters (Langford & Massad, 2014;
183 Pettigrew *et al.*, 2019), based on abundant evidence of freshwater vegetation and mud cracks (Langford
184 & Massad, 2014) and distinct carbonate microfacies of continental origin (Pettigrew *et al.*, 2019). The
185 carbonate microfacies are interpreted to have formed due to microbial processes within either arid or
186 humid aeolian interdunes, saline pans around the edge of desert lakes, and in, or around, an inland
187 evaporitic lake that showed evidence of periodic contraction and expansion (Pettigrew *et al.*, 2019).

188 **METHODS**

189 Ten detailed sedimentary logs were measured through canyons running perpendicular to the strike of
190 the sediments at approximately 3 km intervals along a 15 km long north – south transect. No subsurface
191 core exists for the study area. The logs have a cumulative sediment thickness of 7.5 km and form the
192 basis of the lithofacies within this study. A further two logs (one 60 km to the south of Bluff, and one
193 100 km to the north of the town) were measured to show the spatial variability in sediments preserved,
194 and to provide regional context. The logs were correlated to one another using the top and base of the
195 Cedar Mesa Sandstone Formation as markers, and by tracing prominent units between logs using
196 continuous outcrop.

197 From log 1.4 (Road Canyon), in the middle of the study area (Figs 1 and 2), evaporites were sampled
198 from every evaporitic bed, and from nodules and veins within clastic deposits. The mineralogy of these
199 samples was characterized using X-Ray diffraction (XRD), and triple oxygen and sulphur isotope
200 analyses were conducted upon them. A further ten evaporitic samples were collected from logs 1.4 to
201 1.8 and were cut dry to produce 30 µm thick, unstained thin sections that were subsequently examined
202 for evaporitic textures.

203 Samples for X-Ray Diffraction (XRD) analyses were milled to a fine powder using an agate ball mill.
204 The XRD analyses were performed at the University of St Andrews using a Philips PW1830

205 Diffractometer (Co K α X-ray source), Philips PW1710 Diffractor Control unit, and operating conditions
 206 of 30 kV and 30 mA (Koninklijke Philips N.V., Amsterdam, The Netherlands). Samples were scanned
 207 from 3 to 70° 2 θ using a step size of 0.01° and a counting time of 1 second per step.

208 The extraction and purification of sulphate for stable isotope analysis is described fully in the
 209 Supplementary Information. Measurements of $\delta^{34}\text{S}$ (Table. 3) were made from SO₂ gas using a Thermo
 210 FlashSmart Elemental Analyser (EA) combined with a MAT 253 isotope-ratio mass spectrometer
 211 (IRMS; Thermo Fisher Scientific, Waltham, MA, USA) at the University of St Andrews (Warke *et al.*,
 212 2020). SO₂ was generated by combustion of barite samples (0.14 –0.19 mg) and vanadium pentoxide
 213 powder (1–2 mg) combustion in the presence of O₂ at 1020°C. SO₂ was separated from other gases using
 214 a Gas Chromatography (GC) column after combustion products were carried via helium stream through
 215 tungsten oxide catalysts and copper wire (to reduce excess O₂) and a magnesium perchlorate trap (*ca*
 216 20°C) to remove water. Sample isotopic compositions were calibrated to Vienna Canyon Diablo Troilite
 217 (V-CDT) scale by bracketing barite samples with the international reference standards IAEA-SO5,
 218 IAEA-SO6 and NBS-127, and are reported in the standard delta notation where:

$$219 \quad \delta^{34}\text{S} = 1000 * [((^{34}\text{S}/^{32}\text{S})_{\text{sample}} / (^{34}\text{S}/^{32}\text{S})_{\text{V-CDT}}) - 1] \quad (1)$$

220 Long-term analytical uncertainty is better than 0.3‰ (1 σ).

221 Triple oxygen isotope analysis was performed using a laser fluorination method at the OASIC
 222 laboratory at Louisiana State University (LSU). The protocol is described elsewhere (e.g. Crockford *et*
 223 *al.*, 2018) and is outlined fully in the Supplementary Information. Sample gases were measured with
 224 respect to internal standard LSU-O₂ which is calibrated to Standard Mean Ocean Water (SMOW).
 225 Long-term analytical uncertainty is better than 0.05‰ (1 σ). Measurements of $\delta^{17}\text{O}$ and $\delta^{18}\text{O}$ were made
 226 on purified O₂ using a dual inlet MAT 253 isotope-ratio mass spectrometer (IRMS) and are presented
 227 in the standard delta notation where:

$$228 \quad \delta^{\text{X}}\text{O} = 1000 * [((^{\text{X}}\text{O}/^{16}\text{O})_{\text{sample}} / (^{\text{X}}\text{O}/^{16}\text{O})_{\text{SMOW}}) - 1] \quad (2)$$

229 In Eq. 2, ‘x’ can equal 17 or 18. $\Delta^{17}\text{O}$ values were calculated from $\delta^{17}\text{O}$ and $\delta^{18}\text{O}$ values where:

$$\Delta^{17}\text{O} = 1000 * [\ln(1+\delta^{17}\text{O}/1000) - (0.528 * \ln(1+\delta^{18}\text{O}/1000))] \quad (3)$$

231 As the laser fluorination method causes a mass-dependent fractionation, $\delta^{18}\text{O}$ values were separately
232 evaluated by converting barite to CO gas at 1450°C using a Thermal Conversion Elemental Analyzer
233 (TCEA). Measurements of $\delta^{18}\text{O}$ (Table 3) were made on CO gas using a MAT253 IRMS (isotope ratio
234 mass spectrometer; Thermo Fisher Scientific, Waltham, MA, USA) at the OASIC laboratory at LSU.
235 Long-term analytical uncertainty on $\delta^{18}\text{O}$ is better than 0.5‰ (1 σ) (Peng *et al.*, 2011).

236 **RESULTS**

237 Lithofacies analysis of logged sections (Fig. 2) reveal fifteen lithofacies deposited by subaqueous, sub-
238 aerial and evaporitic processes that form commonly occurring deposits typical of the arid environment.
239 Log data for evaporitic lithofacies are augmented with results of XRD analysis and thin-section
240 observations. The XRD data are summarized in Table 1 and discussed in the text, with all XRD spectra
241 shown in the Supplementary Information (Appendix). The distinctive features of each of the fifteen
242 lithofacies and their interpretations are summarized in Table 2. They have been grouped into nine
243 discrete facies associations based on the dominant depositional process and deposits preserved, each of
244 which is described in detail below.

245 **Aeolian deposits**

246 *Dune associations (AD)*

247 Deposits within the study area are composed of laterally extensive tabular bodies with relatively flat
248 basal bounding surfaces. Lateral extents are over tens of metres, although stratigraphic thicknesses are
249 limited (*ca* 5 m). Deposits of this type are dominantly composed of tabular cross-bedded sandstones
250 (Sxb) with planar foresets (1 to 5 cm in thickness), and rare occurrences of trough cross-bedded
251 sandstones (Stxb) with tangential foresets of similar thickness (1 to 5 cm) (Fig. 3A). The deposits lack
252 large-scale set or coset development, with sets 1 to 5 m and cosets no greater than 10 m in thickness.
253 Planar cross-bedded and trough cross-bedded sets typically interfinger with inversely graded, ballistic-
254 rippled sandstone (Sxr), structureless sandstone (Sm) and gypsum facies (G) along the toesets of the

255 foresets, with small-scale (<50 cm) soft-sediment deformation also a typical feature. The foresets are
256 sporadically disturbed by evaporite deposits in the form of nodules and veins (Fig. 3B).

257 *Interpretation:*

258 Cross-bedded sandstones that display tangential foresets, formed of lamination couplets of structureless
259 to reverse graded sandstone with a thin veneer of finer grained sandstone, represent aeolian dune
260 deposits formed by alternating grainfall and flow processes. This indicates reasonably well developed
261 dune-forms with lee-slope slip-faces close to, or at, the angle of repose (Hunter, 1977; Kocurek, 1981;
262 1991; 1996; Langford & Chan, 1989; Mountney, 2006). Dunes were dominantly straight-crested
263 transverse forms, preserving planar cross-bedded sets with planar set-bounding surfaces. Sporadically,
264 sinuous-crested transverse forms developed, or straight-crested forms broke up or evolved into sinuous-
265 crested bedforms under autogenic processes, due either to changes in migration rate, sediment supply
266 or wind direction (Rubin & Hunter, 1984). The presence of ballistic wind ripples interfingering with
267 the toesets of cross-bedding suggests the presence of dune plinths upon which ballistic ripples
268 dominated. Soft-sediment deformation, typically preserved near the base of the deposits, indicate dunes
269 migrated in the presence of a water table close to, or at, the surface (McKee *et al.*, 1971; Doe & Dott,
270 1980; Horowitz, 1982; Mountney & Thompson, 2002). The presence of gypsum preserved along the
271 base of foresets or toset surfaces, indicates that solute-rich water was drawn up preferential flow
272 pathways in the sediments of the advancing dune as a result of capillary action.

273 ***Sandsheet associations (SH)***

274 Deposits of this type consist of thin, sheet-like, laterally extensive, tabular bodies with flat basal and
275 top surfaces. The deposits have lateral extents over distances greater than tens of metres with limited
276 thicknesses (1 to 2 m), and consist primarily of inversely graded, ballistic-rippled (Sxr) (Fig. 3C and D)
277 or structureless sandstones (Sm), with sporadic crudely developed cross-bedded sandstones (Sxb).
278 Deposits typically grade laterally and vertically into dune and interdune deposits and contain gypsum
279 nodules and veins that deform the bases of the deposits.

280 *Interpretation:*

281 Deposits primarily composed of ballistic-ripple strata (Sxr) or structureless sandstones (Sm) indicate
282 aeolian sandsheets in which limited grain-size variations do not distinguish the bounding surfaces of
283 individual cross-laminated sets (Kocurek, 1981). The sandsheets developed under conditions where
284 sediment supply was insufficient for full-scale dune development (Kocurek & Nielson, 1986; Biswas,
285 2005), resulting in only minor crude cross-bedding. Soft sediment deformation indicates a water table
286 close to the surface, with deformation resulting from water table fluctuations and loading of the
287 sandsheet by the following dune deposits (McKee *et al.*, 1971).

288 **Standing water deposits**

289 *Interdune associations (ID)*

290 Deposits of this type consist of either lensoidal bodies that pinch out over distances of less than one
291 metre, or tabular bodies with larger lateral extents over tens of metres. In each case, thicknesses are no
292 greater than 2 m. Deposits comprise structureless sandstones (Sm), oscillatory current rippled siltstone
293 and sandstone (Swr) and sand-rich carbonates (Lsm). Deposits grade vertically and laterally into dune
294 (AD) and sandsheet (SH) deposits, with sporadic examples occurring in relationship with ponded water
295 (PA) and displacive gypsum (DG) associations. Rhizoliths, bioturbated sediment and mottling are
296 present throughout (Fig. 4C), with sporadic occurrences of convolute-bedded sandstone (Scu),
297 gypsum/anhydrite nodules and efflorescent crusts.

298 *Interpretation:*

299 Associations of this type are interpreted as the deposits of wet and saline interdunes. Carbonate deposits
300 have accumulated in long-lived shallow ponds of standing water above the depositional surface, close
301 to an aeolian dune field that supplied clastic sediment to the interdune (Loope 1981; 1984; Langford &
302 Chan, 1989; Pettigrew *et al.*, 2019). The isolated nature of the deposits signifies laterally restricted
303 interdunes developing in an enclosed or semi-enclosed setting, either as a result of floodwaters trapped
304 between surrounding sinuous-crested dunes, or by rises in the water table in topographical depressions
305 (Loope, 1984; Purvis 1991; Mountney & Jagger, 2004). Oscillatory-ripples are formed from wind shear
306 over the shallow water of the interdune setting (Martell & Gibling, 1991). Rhizoliths and mottling

307 indicate the short-lived presence of water at the surface, probably as a result of fluctuations in the
308 elevation of the water table.

309 ***Ponded waters associations (PA)***

310 Deposits of this type are characterized by laterally extensive (>10 m) tabular bodies with planar basal
311 surfaces, but thicknesses are highly variable, with deposits ranging between 0.2 to 10 m thick. Deposits
312 are composed primarily of sandstones displaying symmetrical oscillatory-ripples (Swr) (Fig. 4B), fine-
313 grained carbonates (Lm) with sporadic rounded gypsum nodules and minor parallel-laminated siltstones
314 (Ssl). Ponded water sediments typically grade vertically and laterally into interdune (ID), unconfined
315 flow (UF) and suspension settlement (SS) deposits. Rhizoliths, bioturbated sediment, mottling and
316 gypsum nodules are typical features. The rhizoliths are horizontal to sub-vertical and the bioturbated
317 sediment consists of thin (<10 cm) branched vertical and horizontal tube-like burrows.

318 ***Interpretation:***

319 A dominance of wave-ripple sandstones and fine-grained carbonates suggests a shallow ponded water
320 setting. Wave-ripples have been generated by wind shear across the water surface (Martell & Gibling,
321 1991) to produce symmetrical ripples. Fine-grained carbonate facies suggest variations in clastic
322 content and prolonged conditions of low-energy standing water (Tucker, 1978; Platt & Wright, 1991
323 Pettigrew *et al.*, 2019) with frequent reworking by wave action. The low diversity of fossil species
324 within the carbonates, coupled with gypsum nodules, suggests an arid, highly saline and probably
325 restricted environment (Cecil, 1990; Flügel, 2004), with the development of rhizoliths indicating
326 surface stabilization around edges of long-standing quiet water bodies (Platt & Wright, 1991; Owen *et*
327 *al.*, 2008).

328 ***Stratified waters/suspension settlement associations (SS)***

329 Deposits of this type comprise laterally extensive tabular bodies with planar basal surfaces. Lateral
330 extents are commonly over tens of metres whereas stratigraphic thicknesses range from 1 m to tens of
331 metres, with the thickest deposits present in the southern portion of the study area (Figs 1C and 8). The
332 deposits are dominantly composed of black to purple coloured siltstones to very fine-grained sandstones

333 (Ssl) (Fig. 4A) intercalated sporadically with oscillatory-rippled sandstones (Swr) (Fig. 4B), fine-
334 grained carbonates (Lm) and horizontally-laminated pedogenic facies (Sfo) (Fig. 4D). The siltstones
335 are typically structureless, mottled, with a sporadically high organic content. Deposits grade laterally
336 into pedogenic facies (Sfo) and typically grade vertically with unconfined flow (UC), or ponded water
337 (PA) deposits.

338 *Interpretation:*

339 Fine-grained siltstones that lack internal sedimentary structures indicate deposition in lakes that were
340 perennial and of greater extent than Ponded Water or Interdune settings, and in which the sediments
341 settled mostly from suspension (Fielding, 1984; Tanner & Lucas, 2007). Sporadic evidence for high
342 organic content within siltstones could indicate depths sufficient enough to cause thermal stratification
343 (Boehrer & Schultze, 2008). Intermittent intercalation with wave-ripples (Swr) shows a shallowing of
344 the water level, potentially caused by climate fluctuations, and an increasing influence from wind shear
345 (Martell & Gibling, 1991). The occurrence of horizontally laminated pedogenic facies suggests
346 stabilization around the margins of a long-standing body of water (Eberth & Miall, 1991).

347 **Flowing water deposits**

348 *Unconfined flow associations (UF)*

349 Deposits of this type consist of thin, sheet-like, laterally extensive bodies with relatively flat to slightly
350 concave-upward, erosional, basal surfaces. Lateral extents are highly variable and range from 1 m to
351 tens of metres, which often branch into thinner sheet-like units, and vertical extents are no greater than
352 1.5 m. This association comprises cross-bedded sandstones (Sfxb) with sporadic mud clasts, arranged
353 into multiple low-angle cross-bedded sets (Fig. 4E) (up to 1 m thick), planar laminations (Sfpl) of 0.5
354 to 1.0 cm thick (Fig.4F) and sporadic climbing-ripple laminated sandstone (Sfrl). Preservation of the
355 full, fining upward succession is rare and the association often grades into suspension settlement
356 siltstones, ponded water deposits or pedogenic facies.

357 *Interpretation:*

358 The flat erosional bases, laterally extensive geometries and sedimentary fill of these associations
359 represents characteristic non-channelized flow incorporating either splay-type or sheetflood-type
360 architectures (Tunbridge, 1981; Sneh, 1983; Stear, 1985; Miall, 1985; North & Davidson, 2012; Priddy
361 *et al.*, 2019; Priddy & Clarke, 2020). A dominance of low-angle cross-bedding and planar-laminated
362 sandstones that stack into fining upward vertical successions indicates a transition from lower-flow
363 regime to upper-flow regime conditions and structures, within high velocity flows that waned quickly
364 (Bridge, 1993; Mitten *et al.*, 2020), whereas thin units that split and branch represent the margins of
365 individual sheetfloods as a result of a rapid reduction in depth associated with an ephemeral, sand-rich
366 sheet flood (Miall 1985; 1996).

367 **Chemically precipitated deposits**

368 ***Bedded gypsum/anhydrite associations (BG)***

369 Deposits of this type have a flat basal surface, and consist of thin, flat bodies. Observed lateral extents
370 are variable, ranging between 0.5 to 5.0 m with vertical extents typically less than 5 m. The deposits
371 consist almost entirely of white to peach (5R 8/3), very fine-grained, crystalline gypsum (G), with
372 outcrop-scale alabastrine and porphyroblastic textures resembling white marble, interbedded with thin
373 beds of siltstones (Ssl). The evaporitic sediments are crudely bedded with sporadically laminated layers
374 of peach coloured elongate gypsum crystals and large botryoidal gypsum nodules (up to 50 cm in
375 height) which bifurcate up from a stem and branch into a cauliflower-like appearance (Fig. 5A). They
376 comprise a matrix of alabastrine gypsum with very minor amounts of anhydrite and quartz. Three
377 distinct textures are recognized: (i) randomly organized angular gypsum (<1 mm) surrounded by brown
378 mudstone clasts, with frequent veins filled with satin spar gypsum (Fig. 5B); (ii) spherical and sub-
379 rounded patches of sand (gypsum and some quartz) (Fig. 5C); and (iii) thin (<1 to 3 mm) undulating
380 beds of bladed gypsum with vertical textures and rare swallowtail crystal morphology (Fig. 5B and C).

381 The bedded gypsum deposits (BG) grade vertically and laterally into suspension settlement (SS),
382 ponded waters (PA) and displacive gypsum (DG) deposits, and occur sporadically in association with
383 unconfined flow (UF) and interdune (ID) deposits.

384 *Interpretation:*

385 These deposits result from the accumulation of primary gypsum and anhydrite by precipitation in
386 shallow, saline waters (Kendall, 1981; Handford, 1991; Warren, 1991) where modern studies indicate
387 that the three textures observed are typically present together (Benison *et al.*, 2007). Laminated layers
388 with elongate crystals are indicative of selenite, signifying bottom-growth of crystals within standing
389 brines (Benison & Goldstein, 2000, Benison & Goldstein, 2001). Selenite has been dehydrated to
390 anhydrite during burial diagenesis and rehydrated to secondary gypsum during exhumation (Gundogan
391 *et al.*, 2005). Siliciclastic sediment intercalated with the bedded evaporite crystals indicates coeval
392 clastic deposition by either fluvial, lacustrine or aeolian processes (Foster *et al.*, 2014).

393 ***Displacive gypsum/anhydrite associations (DG)***

394 These deposits consist of thin undulating bands of laterally extensive bodies with undulating basal and
395 top bounding surfaces. Lateral extents range from 1 m to tens of metres and thicknesses range from 0.5
396 to 2.0 m. The deposits are dominantly composed of crystalline gypsum (G) with alabastrine and
397 porphyroblastic textures, and minor amounts of anhydrite, celestine and quartz (Table 1) (Fig 6B to D).
398 The evaporites comprise a matrix of interlocking alabastrine crystals (<1 mm) with randomly orientated
399 large (1.5 to 4.0 mm) cubic to sub-cubic halite crystals filled with a clear cement with bladed cement
400 crystals present along the edges (Fig. 6E). Small gypsum nodules (up to 20 cm in thickness) and
401 laminated-bands of enterolithic convoluted folds (Fig. 6C and D), polygonal hummocks (Fig. 6A) and
402 chicken-wire structures (Fig. 6B) are common features within these deposits. The nodules are
403 interbedded either with thin beds (up to 20 cm thick) of pastel blue, very fine to fine-grained, parallel-
404 laminated gypsum-bound sandstone (Gspl) within a gypsiferous matrix and cement, or with very fine
405 to fine-grained pedogenic (Sfr) facies with frequent gypsum nodules and veins.

406 The displacive gypsum (DG) deposits grade vertically and laterally into bedded gypsum (BG)
407 suspension settlement (SS), ponded waters (PA) deposits, and typically occur in association with
408 unconfined flow (UF) and interdune (ID) deposits.

409 *Interpretation:*

410 Interbedded gypsum-bound sandstone with intercalated gypsum nodules and veins relate to saline pan
411 and saline mudflat deposition along the margins of saline lakes (Lowenstein & Hardie, 1985).
412 Enterolithic and chicken-wire structures indicate very early diagenetic (syndimentary) displacive
413 growth of anhydrite nodules which were partly or completely hydrated to secondary gypsum under
414 post-depositional conditions (i.e. early and/or late diagenesis) and the exhumation process (Butler,
415 1970). Polygonal hummocks form tepee structures due to desiccation of saline waters and fracturing of
416 salt crusts into polygonal shapes (Warren, 2016), the presence and preservation of such structures
417 suggests a calm, low-energy environment (Lokier & Steuber 2008; 2009). The matrix textures (Fig. 6E)
418 suggest that the host sediment that was originally mud (potentially gypsum-rich mud) with early
419 diagenetic displacive halite crystals. The largest halite crystals (Fig. 6E) show euhedral cubic to sub-
420 cubic habits and are randomly oriented within the finer matrix. The displacive halite crystals appear to
421 have been dissolved (Casas & Lowenstein, 1989; Benison *et al.*, 2015), forming crystal moulds that
422 were later filled with anhydrite or gypsum bladed cement crystals along the edges. The remainder of
423 the crystal mould has been filled with a clear halite cement. The interlocking crystal mosaic present
424 within the fine matrix indicates that the original mud underwent neomorphism.

425 ***Brecciated gypsum/anhydrite associations (BrG)***

426 These deposits consist of thick laterally extensive (up to 10 m) lensoidal bodies with flat to undulating
427 basal surfaces. Observed thicknesses range from 1 to 5 m and deposits of this type are especially
428 prevalent at the top of the formation. The deposits are composed almost entirely of weathered mounds
429 (Fig. 7A and B) of white to peach, very fine-grained brecciated gypsum (G) with minor anhydrite (Table
430 1), and minor components of arid pedogenic facies (Sfr), carbonates (Lm) and gypsum-bound sandstone
431 (Gspl). The matrix of the gypsum deposits is composed of elongate rectangular selenite crystals with
432 rounded ends (1 to 10 mm in horizontal thickness) (Fig.7C). Carbonate deposits are thin (<40 cm) and
433 show wavy laminations, typically interbedded with thin (<20 cm) deposits of gypsum-bound sandstone
434 (Gspl). Pedogenic facies are also typically thin (<50 cm) with frequent gypsum nodules and veins.
435 Deposits of this type (BrG) occur in vertical and lateral relationship with dune (AD) and sandsheet (SH)
436 deposits.

437 *Interpretation:*

438 Brecciated gypsum deposits are interpreted as gypsum dunes resulting from the deflation of gypsum
439 sediment formed in lacustrine and playa lakes during more humid climatic conditions (Szyrkiewicz *et al.*
440 *al.*, 2010). During the onset of more arid conditions, desiccation of these playa lakes led to the formation
441 of microbial mats, present in the form of carbonate laminations (Pettigrew *et al.*, 2019). Aeolian
442 reworking of desiccated lake beds produced gypsiferous dune deposits controlled by a near-surface
443 groundwater table (Szyrkiewicz *et al.*, 2010) and arid soils (Sfr) around lake edges (e.g. Lawton &
444 Buck, 2006). The weathered appearance of the gypsum dunes results from their poor consolidation
445 compared to their clastic counterparts and is a post-depositional burial, dissolution and compaction that
446 destroyed primary bedding fabrics, (Fenton *et al.*, 2014). The dune sediment matrix shows
447 characteristics of aeolian deposits of gypsum/anhydrite that have been slightly altered by diagenesis
448 (hydration–dehydration reactions) and that compare with morphologies of modern gypsum sand dunes
449 and sand flats, such as those at White Sands in New Mexico, in the Rio Grande Rift (Wilkins & Currey,
450 1997; Anderson *et al.*, 2002; Langford, 2003; Kocurek *et al.*, 2007; Szyrkiewicz *et al.*, 2010) and the
451 Olympia Undae Dune Field (Mars) (Langevin *et al.*, 2005; Fishbaugh *et al.*, 2007).

452 **Spatial distribution of deposits**

453 Spatial analysis of the deposits indicates a general increase in the relative proportion of ‘water derived’
454 deposits southward over the study area (Fig. 8). In the north, logs 1.0 to 1.3 (covering a transect
455 approximately 60 km long from Canyonlands to the south of US95) demonstrate dominantly aeolian
456 deposits of dune and sandsheet, with subordinate interdunes, and indicate a dry to sporadically damp
457 setting throughout times when deposition of the Cedar Mesa was controlled by aeolian processes. South
458 of US163, log 1.9 demonstrates dominantly lacustrine deposits of suspension settlement and unconfined
459 flow associations and indicates a perennial lake throughout Cedar Mesa times. From US163 northward
460 for approximately 30 km (logs 1.4 to 1.8; the central part of the studied area), the Cedar Mesa contains
461 aeolian and standing water deposits mixed with chemically precipitated and flowing water deposits in
462 varying proportions, showing the interbedded and competing nature of the aeolian and lacustrine
463 settings.

464 **DEPOSITIONAL SETTINGS**

465 **Dry and saline sandflats**

466 In the north of the study area between Canyonlands and Blanding (logs 1.0 to 1.3) (Fig. 2), thick
467 accumulations of dune associations (AD) and laterally extensive sandsheet (SS) deposits dominate the
468 successions, with minimal accumulations of evaporitic associations. Whereas, in the rest of the study
469 area, between Blanding and Bluff (logs 1.4 to 1.8), the deposits of aeolian dunes (AD) are small,
470 isolated, typically evaporite-rich, haloturbated and interbedded with evaporate-rich interdunes (ID) and
471 arid palaeosols that typically display evaporite nodules, gypsum veins and efflorescent crusts.

472 A dominance of dune and sandsheet associations in the north (Canyonlands to Blanding) suggests
473 deposition in a dry sandflat setting most probably at the edge of the main Cedar Mesa erg. Dune deposits
474 lack salt influence and are comparatively larger than those in the south with trough cross-bedding
475 arranged into multiple sets and cosets. Limited interdune and sandsheet deposits are preserved between
476 dune sediments (Mountney & Jagger, 2004; Vackiner *et al.*, 2011; Antrett, 2013).

477 However, between Blanding and Bluff, the deposits represent a saline sandflat environment, sometimes
478 referred to as a wet aeolian or saline aeolian sabkha (Warren, 2016). Small isolated dunes migrated over
479 saline saturated ground waters, with a saline water table in contact with, or above, the depositional
480 surface (Driese, 1985; Warren, 2016; Zuchuat *et al.*, 2019). The deposits of the saline sandflats show
481 varying degrees of salt preservation, however haloturbation is typical, even where little capillary salt is
482 preserved (Warren, 2016). Overprinting by extensive disruptive evaporite crystal growth can modify or
483 destroy primary depositional features (Ahlbrandt & Fryberger, 1981; Warren, 2016) which typically
484 consist of subaqueous current and wave ripples, wavy and contorted bedding, adhesion structures,
485 bioturbation, desiccation cracks, megapolygons (Warren & Kendall, 1985), deflation surfaces and
486 wind-ripple lamination (Martin & Evans, 1988; Goodall *et al.*, 2000; Mountney & Thompson, 2002;
487 Warren, 2016).

488 Wet interdune/arid outwash plain

489 This setting is characterized by thick deposits of ponded water (PA) and interdune (ID) associations
490 interbedded with minor occurrences of unconfined flow (UF), stratified waters (SS) and dune
491 associations (AD). Dune deposits are typically straight crested (Sxb) and interbedded laterally and
492 vertically with interdune deposits of massive sandstone (Sm), oscillatory current rippled sandstone
493 (Swr) and rooted palaeosols (Sfr).

494 These deposits represent a wet interdune/arid outwash plain. Sporadic occurrences of dune deposits
495 separated from one another by wide interconnected wet interdunes suggests a setting with a degree of
496 aridity but limited aeolian sediment supply, such as at the edge of larger dune field. These areas typically
497 have higher water tables during humid periods, forming ponded water deposits and carbonates
498 (Pettigrew *et al.*, 2019). Occasional higher energy events input clastic material through unconfined
499 flows to interdune areas reworking sediment (Howell & Mountney, 1997; Mountney, 2006).

500 Perennial lake

501 This setting is characterized by thick stratified water associations (SS) primarily comprising deposits of
502 siltstone (Ssl) interbedded with lacustrine carbonates (Lm and Lsm), oscillatory current rippled
503 sandstone (Swr) and palaeosols (Sfo) with sporadic unconfined flows (UF).

504 These deposits thicken towards the south and indicate a regionally-extensive low-energy, shallow
505 lacustrine setting influenced by the sporadic input of sand-rich water from unconfined surface run off.
506 Organic rich siltstones indicate relative longevity and sufficient depths of a perennial lake for thermal
507 stratification. Wave ripples suggest periods of lake level fluctuations as shallowing water levels
508 increased the influence of wind shear and/or wave action reworking along the margins of the lake
509 (Bridge & Demicco, 2008). The lake margin deposits grade into humid palaeosols indicating
510 stabilization around the lake edge.

511 Ephemeral saline lake-saline pan/mudflat

512 The deposits of ephemeral saline lakes are characterized by bedded gypsum (BG) with minor, laterally
513 restricted deposits of the ponded water (PA) associations which grade into displacive (DG) and
514 brecciated gypsum (BrG) associations of a saline pan/mudflat.

515 This setting represents saline pan and mudflat environments with ephemeral saline lakes which formed
516 during arid periods via the evaporation and desiccation of previously long-lived lakes. In the main
517 submerged areas of the lake, during desiccation (Lowenstein & Hardie, 1985) the earliest precipitates
518 nucleate as thin rafts of connected platy euhedral cumulate halite crystals at the water–air interface that
519 are held by surface tension until they grow too large and sink (Schreiber & Kinsman, 1975; Castens-
520 Seidell, 1984; Last, 1984; Alderman, 1985; Hardie *et al.*, 1985; Lowenstein & Hardie, 1985; Smoot &
521 Lowenstein, 1991). As evaporation of the lake progresses, salts are concentrated in the remaining water
522 to form smaller ephemeral saline lakes in which sulphate and calcium concentrate in shallow waters.
523 This facilitates the precipitation of bottom-nucleating, vertically-elongated gypsum crystals (e.g.
524 Kendal, 1978; Last, 1984; Smoot & Lowenstein, 1991). The fibrous/prismatic gypsum grows
525 perpendicularly to the substrate in contact with the brine (bottom-growth) (Kendall & Harwood, 1996).
526 These crystals increase in size and gradually became a mosaic, formed by displacive growth in very
527 shallow water and/or in the capillary fringe-groundwater zone, as a result of either a fall in water levels
528 or subaerial exposure (e.g. Warren, 2006). Complete desiccation of the saline lake produces polygonal
529 fractures in the salt pan that typically fill with detrital evaporitic or clastic sediment. They preserve as
530 efflorescent crusts (Smooth & Castens-Seidell, 1994) with crack fills, or so-called ‘tepee structures’
531 where the fill incorporates displacive evaporites (Warren, 1983; Lowenstein & Hardie, 1985; Lokier &
532 Steuber, 2009).

533 Saline mudflats form in dominantly emergent settings and are characterized by fine-grained clastic
534 sediments, typically clay minerals such as montmorillonite and illite (Brooks & Ferrell, 1970). These
535 sediments are deposited by settlement following influxes of sediment-laden floodwater into shallow
536 ephemeral lakes. With lake contraction, the sediments are exposed. Saline mudflats have little (if any)
537 vegetation and have a saline groundwater table only centimetres to tens of centimetres below the surface

538 which may fluctuate up and down in response to climate. Variations in ground water levels and
539 circulation promote subsurface phreatic evaporite growth as random crystals or nodules, or as
540 concentrated thick layers, displacing surrounding sediment to form chickenwire and enterolithic
541 structures (Smoot & Lowenstein, 1991; Boggs & Boggs, 2009; Warren, 2016). On the surface of the
542 mudflats, precipitation from the evaporation of saline ground waters forms efflorescent evaporites,
543 either as powdery undulating surfaces or as hard crystalline crusts (Smoot & Lowenstein, 1991; Warren,
544 2016). Subaerial exposure and desiccation of gypsum in the vadose zone results in fractures and the
545 reworking of fragments by water inflow (Kendall & Harwood, 1996; Abrantes Jr *et al.*, 2016).

546 **SUMMARY DEPOSITIONAL MODELS**

547 Two models showing the intermediate environments between a desert erg and desert lake have been
548 distilled from the depositional settings. Within each model, clear variations in sedimentology and
549 characteristics of more humid or more arid conditions can be recognized (Fig. 9).

550 **Dune-dominant, erg-marginal system**

551 During times of relatively higher humidity, the erg-marginal system is characterized by a wet, marginal,
552 aeolian erg interacting with unconfined run off (e.g. Mountney & Jagger, 2004) that feeds small saline
553 lakes and interdunal areas (Fig. 9). Evaporite development was limited, with sufficient surface water
554 for vegetation and soil formation.

555 Throughout subsequent arid periods accompanied by a drop on the groundwater table, aeolian dunes
556 increased in magnitude and frequency as they became contiguous with the central erg. They formed
557 dune fields, dominated by straight-crested dunes evolving through time into sinuous-crested forms with
558 well-developed dune plinths, within dry sandflats. The facies belt graded distally into a saline sandflat
559 setting, in which small-scale, sinuous-crested, aeolian dunes migrated over saline-rich damp or wet
560 interdunal areas. Locally the solute-rich substrate allowed for the direct precipitation of evaporites and
561 carbonates, intercalated with aeolian calibre clastic sediments, to form saline interdune deposits.
562 However, evidence for the evaporitic nature of the system as a whole is limited primarily to displacive

563 gypsiferous nodules and haloturbation in aeolian bedforms. A wet solute-rich substrate severely limited
564 sediment supply to dunes and consequently led to isolated barchanoid duneforms with laterally
565 interconnected saline interdunes.

566 Although water played a role in the formation of sediments within the erg-margin, the dominant
567 sediment transport processes are wind-driven. Fluctuations in the availability of water at the surface
568 during arid-humid cycles probably resulted from oscillations in the water table, rather than from
569 sustained surface flow from outside the system that fed significant and long-standing bodies of water.

570 **Lake-marginal system**

571 At a time of relatively higher humidity, the lacustrine-marginal system is characterized by the
572 dominance of lacustrine deposits (Fig. 9). A thick lacustrine succession, which predominantly settled
573 from suspension, is interbedded with either surface run off or carbonate sediments depending on the
574 rate of water input and the magnitude of its clastic load. Around the lake margins, any aeolian dunes
575 were small-scale, barchanoid and strongly isolated between sheetflood-like deposits of fluvial origin
576 and vegetated overbank.

577 During subsequent periods of relatively higher aridity, the lake contracted until extensive saline pans
578 and mudflats developed over the edges of the lacustrine depression, associated with the development of
579 localized ephemeral saline lakes. In the lake, the lack of clastic input from unconfined surface run off
580 during periods of increased aridity is coupled with the concentration of dissolved salts as the water
581 evaporated. This led to the deposition of evaporitic sediments dominated first by halite rafts, formed on
582 the lake surface, and later by bottom nucleating gypsum. Enterolithic and displacive salts grew around
583 the edges of the evaporitic lake due to capillary and phreatic growth from the infiltration of saline
584 groundwater, forming pervasive saline mudflats. Finally, subaerial exposure and desiccation of saline
585 cumulates led to the formation of polygonal growth structures coupled with reworking of cumulates by
586 wind-action into aeolian-gypsum dunes. Aeolian dunes deposits are limited despite the increased
587 aridity. Surface and enterolithic mineral growth in these areas severely limits the exposed sediment that

588 is available for aeolian transport, and traps much wind-blown sediment on damp surfaces.
589 Consequently, the dune field shows little to no growth as the lacustrine environment dried out.

590 **SULPHUR ($\Delta^{34}\text{S}$) AND TRIPLE OXYGEN ($\Delta^{17}\text{O}$) ISOTOPE**

591 **RESULTS**

592 In Log 1.4, measured sulphate $\delta^{34}\text{S}$ values range from $+13.4 \pm 0.3\text{‰}$ to $+14.5 \pm 0.3\text{‰}$, $\delta^{18}\text{O}$ values range
593 from $+12.8 \pm 0.5\text{‰}$ to $+16.69 \pm 0.5\text{‰}$, and values of $\Delta^{17}\text{O}$ range from $-0.27 \pm 0.05\text{‰}$ to $-0.06 \pm 0.05\text{‰}$
594 (Table 3, Fig. 10). There is no clear relationship between gypsum facies and either $\delta^{34}\text{S}$, $\delta^{18}\text{O}$ or $\Delta^{17}\text{O}$
595 values. In the case of $\Delta^{17}\text{O}$ values, some of the most negative values identified occur in veins. There is
596 no clear stratigraphic variation in $\Delta^{17}\text{O}$ values (Fig. 10). In the lower portion of Log 1.4 (approximately
597 0 to 20 m; $n = 5$) $\delta^{34}\text{S}$ and $\Delta^{17}\text{O}$ values co-vary with less negative $\Delta^{17}\text{O}$ values corresponding to less
598 negative $\delta^{34}\text{S}$ values (Fig. 10). Throughout the rest of the section, however, there is no discernible
599 covariation between $\delta^{34}\text{S}$ and $\Delta^{17}\text{O}$ values.

600 Stratigraphic variation in $\delta^{34}\text{S}$ can be only tentatively identified given the *ca* 1‰ range in the dataset,
601 which is small in comparison to the error margins on each measurement. However, two shifts from
602 relatively larger to smaller $\delta^{34}\text{S}$ values can be visually identified between 0 m and 55 m ,and also
603 between 55 m and 104 m (Fig. 10). These upward trends in $\delta^{34}\text{S}$ values are matched by similar trends
604 in $\delta^{18}\text{O}$ values which decrease gradually over the same intervals (Fig. 10).

605 **Interpretation of isotopes**

606 A coupled and gradual upward decrease in $\delta^{34}\text{S}$ and $\delta^{18}\text{O}$ values suggests the existence of at least two
607 evaporative cycles within the logged section, with the decreasing values reflecting Rayleigh
608 fractionation effects within a closed system (Raab & Spiro, 1996). In such a scenario it seems likely
609 that some freshening of the closed system occurred at *ca* 55 m. Above *ca* 55 m, $\delta^{34}\text{S}$ and $\delta^{18}\text{O}$ values
610 return to those seen at *ca* 0 m (Fig 10). These trends, however, can only be tentatively identified as the
611 $\delta^{34}\text{S}$ values at the base and top of the lower cycle (0 to 55 m) are within error of one another (Fig. 10).
612 There is more distinction in the upper cycle, but the range of $\delta^{34}\text{S}$ values remains $<1.1\text{‰}$ (Fig. 10). The

613 $\delta^{18}\text{O}$ profile of Log 1.4 mirrors the trends in $\delta^{34}\text{S}$ but values at the top and bottom of each cycle are,
614 considering error, distinct from one another.

615 The mean $\delta^{34}\text{S}$ value of 14.0‰ (n = 28 ; $1\sigma = 0.3$) is broadly in accordance with the marine $\delta^{34}\text{S}$ curve
616 during the early Permian, or late Carboniferous periods (Fig. 11A; Claypool *et al.*, 1980; Kampschulte
617 & Strauss; 2004). Therefore, the simplest interpretation of the $\delta^{34}\text{S}$ values in Log 1.4 is that they
618 represent precipitation of gypsum from marine waters. Such an interpretation has been favoured by
619 previous workers (Stanescu & Campbell, 1989).

620 It is important to note that mass-independently fractionated isotopic signatures, like $\Delta^{17}\text{O}$ within
621 sedimentary sulphate, are not affected by mass-dependent processes – like evaporation – unless the
622 sulphate anion complex is disturbed, for example in microbial sulphate reduction (Bao *et al.*, 2015). In
623 such cases, any non-zero $\Delta^{17}\text{O}$ value would be erased. Although the range of $\Delta^{17}\text{O}$ values is
624 comparatively small (*ca* 0.21‰) with respect to the range of $\delta^{34}\text{S}$ values, such values are typical of
625 evaporitic sulphates that preserve an atmospheric signal inherited from tropospheric O_2 and imparted to
626 sulphate during sulphide oxidation. For the purposes of comparison, modern tropospheric O_2 has a $\Delta^{17}\text{O}$
627 of -0.47‰ (Pack *et al.*, 2017) but the authors would expect only a fraction of that signal (i.e. *ca* 21 to
628 34%) to be transferred to sulphate during sulphide oxidation (Kohl & Bao, 2011). The only processes
629 known to cause mass-independent fractionation of oxygen isotopes and generate negative $\Delta^{17}\text{O}$ values
630 in atmospheric O_2 are ozone creation and destruction reactions (Thiemens, 2006). The $\Delta^{17}\text{O}$ values
631 measured here, which are as low as -0.27‰, show an unambiguous atmospheric signature.

632 There are relatively few measurements of Phanerozoic $\Delta^{17}\text{O}$ values, but most studies to date have
633 focused on Cambrian and Permian marine and non-marine evaporites (Bao, 2005; Bao *et al.*, 2008)
634 reflecting the uneven distribution of evaporites over time (Warren, 2016). Published Permian $\Delta^{17}\text{O}$
635 values range anywhere from +0.02 to -0.16‰ (n = 26) and are restricted to the mid to late Permian
636 Period (Bao, 2005; Bao *et al.*, 2008); there are no noted measurements from the early Permian Period.

637 Hence, the negative $\Delta^{17}\text{O}$ values measured in the Cedar Mesa Sandstone Formation: (i) show an
638 unambiguous atmospheric fingerprint within sulphate; and (ii) are some of the most negative $\Delta^{17}\text{O}$

639 values measured in the Permian Period, and in the Phanerozoic Eon in general. The measurements
640 herein, therefore, could provide important new insight into the temporal variability of $\Delta^{17}\text{O}$ over the
641 Phanerozoic. However, it is important that the source and age of the sulphate complex analysed is
642 accurately known with respect to the age of the sulphate-hosting formation, as discussed below.

643 **DISCUSSION**

644 Lithofacies analysis of outcrop, augmented by thin section and XRD analyses, suggest a fully
645 continental setting for the evaporite-rich sediments of the Cedar Mesa sandstone in the vicinity of Bluff,
646 Utah. Deposition occurred within a transitional zone between a well-established aeolian erg and a saline
647 lake, with the dominance of one depositional setting over the other controlled by variations in aridity.
648 A continental interpretation for the deposits is further supported by the recent published work of
649 Langford & Massad (2015) and Pettigrew *et al.* (2019).

650 However, isotopic analyses of evaporitic samples collected for this study suggest a marine influence
651 upon deposition of these sediments. The range of measured $\delta^{34}\text{S}$ values (+13.4 to +14.5‰) fall within
652 error of late Carboniferous to early Permian marine values (i.e. $+14.6 \pm 0.7\%$; Kampschulte & Strauss,
653 2004) despite a distinct lack of lithofacies relationships and fossil content that typify successions
654 affected by marine influence or seawater incursions. This apparent conflict between the geochemical
655 and sedimentological data of this study, and the implications for interpreting erg-marginal evaporitic
656 depositional settings in general, are discussed in following sections.

657 **Isotopic signatures: continental or marine?**

658 Interpreting the isotopic signatures within evaporites is typically problematical due to recycling of
659 marine signatures (see Taberner *et al.*, 2000), lack of petrographic analyses to assess the impact of
660 diagenesis, and the use of isotopic values from continental gypsum deposits to construct seawater curves
661 (see Denison *et al.*, 1998). The presence of halite and/or gypsum is not an indicator of marine
662 environments or parent waters, although some workers now use those same curves to attribute a marine
663 origin to halite/gypsum. There is a wide range of $\delta^{34}\text{S}$ and strontium isotope ($^{87}\text{Sr}/^{86}\text{Sr}$) values within

664 saline lakes, with values that overlap with seawater signatures. Benison & Bowen (2013) reported $\delta^{34}\text{S}$
665 values from 17.0 to 20.4 within modern saline lakes in Western Australia, compared to modern seawater
666 $\delta^{34}\text{S}$ values of $20.3(\pm 0.3)\text{‰}$. Therefore, although $\delta^{34}\text{S}$ values derived for this study overlap with the
667 marine curve, this does not necessarily imply that they precipitated from seawater. It is, therefore,
668 inaccurate to use only the isotopes from evaporite minerals to infer marine influence and interpret
669 ancient seawater chemistry if it cannot independently demonstrate (for example, through macro-facies
670 and micro-facies analysis) that the depositional setting has experienced marine incursions.

671 Another possibility that explains the observed $\delta^{34}\text{S}$ values equally well is that the sulphate in the gypsum
672 facies has been recycled from an older marine evaporite deposit (Palmer *et al.*, 2004). This requires that
673 there is a suitable unit in the hinterland. In the case of the Cedar Mesa Sandstone, this may be the late
674 Pennsylvanian (*ca* 311 to 306 Ma) Paradox Formation (Trudgill, 2011) which contains abundant
675 deposits of marine evaporites (Nuccio & Condon, 1996) and may have been exposed during deposition
676 of the Cedar Mesa.

677 Evaporites within the Paradox Formation possess a $\delta^{34}\text{S}$ value of $+13.3\text{‰}$ ($n = 1$; Claypool *et al.*, 1980).
678 Other studies, with larger datasets, have estimated that similarly aged marine evaporites would be likely
679 to have $\delta^{34}\text{S}$ values ranging from $+14.6(\pm 0.7)$ to $+15.1(\pm 1.7)\text{‰}$ ($n = 26$; Kampschulte & Strauss, 2004).
680 The mean $\delta^{34}\text{S}$ value measured in this study, $+14.0\text{‰}$ ($1\sigma = 0.4\text{‰}$; $n = 28$), lies between the two and
681 within error of both late Carboniferous and early Permian evaporites and marine evaporites of similar
682 age to the Paradox Formation (Kampschulte & Strauss, 2004) (Fig. 11A).

683 Thus, it is not sufficient to argue, on the basis of ambiguous sulphur isotope values, that evaporites in
684 the Cedar Mesa Sandstone Formation were deposited following marine incursions as has been
685 previously proposed (Stanescio & Campbell, 1989). There are no macro-facies or micro-facies evidence
686 supporting marine influence, and significant evidence arguing against it. The most parsimonious
687 explanation, that reconciles the sedimentological, stratigraphic and isotopic evidence, is that the
688 sulphate within the early Permian Cedar Mesa Sandstone Formation was derived from the dissolution
689 of the older marine evaporites of the late Carboniferous Paradox Formation. Additionally, the recycling

690 of sulphate from the older Paradox Formation may explain why $\delta^{34}\text{S}$ and $\Delta^{17}\text{O}$ values in the Cedar Mesa
691 Sandstone Formation show no systematic stratigraphic trend or relationship with climatic condition
692 and/or facies.

693 **Controls on deposition, tectonics versus climate**

694 In the absence of marine input to explain the presence of evaporitic strata of the Cedar Mesa Sandstone,
695 other controls upon their deposition should be considered. Tectonics and climate are the primary
696 controls on continental depositional systems (Vail *et al.*, 1991; Quigley *et al.*, 2007), controlling both
697 the type and amount of sediment supplied to the basin and the rate at which accommodation is created
698 and filled (Vail *et al.*, 1991). Tectonic activity creates elevated landscapes, as well as generating
699 structural lows and depocentres, while climate influences the discharge and sediment availability.
700 Tectonic and thermal subsidence are recognized as primary controls on sabkha/playa lake formation by
701 generating tectonic sagging in which water can pool and subsequently evaporate (e.g. Mertz & Hubert,
702 1990).

703 However, deposition of the Cedar Mesa Sandstone Formation coincided with a reduction in the
704 subsidence rate of the Paradox Basin, and thus an accompanying reduction in the rate at which
705 accommodation space was generated. A maximum of approximately 2.7 km of sediment was deposited
706 at an estimated sedimentation rate 84 m/Myr by the end of the Pennsylvanian Epoch (Huntoon *et al.*,
707 1996; Nuccio & Condon, 1996). By contrast, approximately 1.8 km of sediment was deposited at a rate
708 of 40 m/Myr during the Permian Period that overfilled the basin. (Huntoon *et al.*, 1996; Nuccio &
709 Condon, 1996; Condon, 1997; Barbeau, 2003). The overfilled state resulted in a basinward progradation
710 of facies belts due to limited accommodation space (Mountney & Jagger, 2004). This may indicate that
711 basin-scale tectonics had negligible effect on the preserved deposits, acting only as a means of
712 generating accommodation through a relatively constant subsidence rate.

713 Glacio-eustatic processes have been recognized as a driving mechanism for the controls on the central
714 and marginal erg deposits of the Cedar Mesa Sandstone (Mountney & Jagger, 2004) where twelve
715 separate erg sequences can be identified within the aeolian sediments (Mountney, 2006). The sequences

716 reflect 412 kyr Milankovitch eccentricity cycles, which forced cyclic changes in climate (Mountney,
717 2006). It follows that the Cedar Mesa Sandstone was probably influenced by periods of humid–arid
718 climate variations during the early Permian Period, which resulted in periods of contraction and
719 expansion within the lacustrine system in the south of the study area (Figs 8 and 9). Arid climatic
720 intervals led to the development of saline mudflats associated with ephemerally saturated carbonate
721 ponds and saline pans in the south of the study area, and the development of aeolian features
722 predominantly in the north. The presence of diagenetic halite (Fig. 6E) is an indirect proxy for saltier
723 and drier conditions than have previously been attributed to the depositional environment and
724 palaeoclimate of the Cedar Mesa, and is consistent with age-equivalent rocks elsewhere in western
725 equatorial Pangea (Parrish, 1993; Kessler *et al.*, 2001; Ricardi-Branco, 2008; Tabour *et al.*, 2013;
726 Falcon-Lang *et al.*, 2015).

727 Despite a probable dominance of climatic controls upon deposition of sediments of the Cedar Mesa
728 Sandstone, the influence of localized structure upon the position of evaporite deposits within the basin
729 cannot be fully discounted. Large-scale and deep-seated faults of Pre-Cambrian age that bound the study
730 area, and that have been purported to be active during Permian times (Kelley, 1955; Huntoon, 1993;
731 Davis, 1999; Mynatt *et al.*, 2009; Hilley *et al.*, 2010), may have produced localized lows into which
732 surface water was focused to produce wetter conditions, even during arid climatic phases, in which erg-
733 marginal evaporitic environments developed (Pettigrew *et al.*, 2019). It is also not beyond the realms
734 of imagination to suggest that displacement on these faults may have brought older marine
735 Carboniferous evaporites into contact with the water table, or to surface (Fig. 11), and that this scenario,
736 coupled with diapiric exposures of evaporitic sediments in the proximal basin (Trudgill, 2011; Venus
737 *et al.*, 2015), may have contributed to the recycling of the isotopic signature. At present however, the
738 authors have no direct geological evidence to support this hypothesis.

739 **Implications for using $\Delta^{17}\text{O}$ as a palaeoatmospheric proxy in non-marine evaporites**

740 The presence of a dominant late Carboniferous to early Permian marine $\delta^{34}\text{S}$ signal in the gypsum
741 deposits analysed here, coupled with the sedimentological constraints precluding marine influence,
742 suggests that the sulphate incorporated into Cedar Mesa was recycled from an older marine evaporite

743 succession. This is entirely reasonable given that the global fluvial sulphate flux is known to be derived
744 from pyrite oxidation and dissolution of calcium sulphate minerals (Halevy *et al.*, 2012; Wortmann &
745 Paytan., 2012). The geology of the hinterland exerts a strong influence on the isotopic composition of
746 fluvial sulphate (Burke *et al.*, 2018; Waldbeck *et al.*, 2019). Given that the Paradox Formation may
747 have been uplifted and possibly even exposed in the footwalls of faults during the Permian Period
748 (Kelley, 1955; Huntoon, 1993; Davis, 1999; Mynatt *et al.*, 2009; Hilley *et al.*, 2010; Pettigrew *et al.*,
749 2019), it seems probable that the hinterland and local area contained a significant proportion of late
750 Carboniferous evaporites. The dissolution of these evaporites, and their incorporation into short-lived,
751 short-distance (tens of kilometres) fluvial systems in the Paradox Basin, may have led to a fluvial
752 sulphate flux that was dominated by sulphate recycled from older evaporite deposits.

753 However, $\delta^{34}\text{S}$ and $\Delta^{17}\text{O}$ values in the Cedar Mesa Sandstone may not only record recycling of older
754 evaporite signals. Contemporaneous sulphide oxidation in other areas of the drainage basin may also
755 have delivered sulphate to the depositional environments via fluvial systems. It might be expected that
756 this sulphate flux would have a less positive (fluvial) $\delta^{34}\text{S}$ value (Burke *et al.*, 2018) and a more negative
757 $\Delta^{17}\text{O}$ value (Waldbeck *et al.*, 2019), producing coupled $\delta^{34}\text{S}$ and $\Delta^{17}\text{O}$ trends; however, with the
758 exception of the lowermost portion of Log 1.4, this is not evident. Additionally, $\delta^{34}\text{S}$ values in the Cedar
759 Mesa Sandstone are consistent with late Carboniferous–early Permian seawater $\delta^{34}\text{S}$ values. If anything,
760 the average $\delta^{34}\text{S}$ value in the Cedar Mesa Sandstone (+14.0‰) is marginally higher than the known
761 Paradox Formation value (+13.3‰; Claypool *et al.*, 1980). It is likely that significant mixing of a
762 sulphide oxidation flux into the fluvial sulphate would have lowered the bulk-rock $\delta^{34}\text{S}$ value.
763 Therefore, this process is considered to be a fairly insignificant sulphate source in comparison to the
764 recycling of older marine evaporites. This is in line with other studies that have found limited evidence
765 for alteration of evaporite-derived bulk rock $\delta^{34}\text{S}$ values by oxidative sulphide weathering in this study
766 area (Breit *et al.*, 1990).

767 The findings of this study hold very important implications for the interpretation of $\Delta^{17}\text{O}$ in continental
768 sulphates, as they too must also have been recycled. As discussed above, continental evaporites could
769 represent a better target for studies attempting to reconstruct coupled biosphere–atmosphere interaction

770 using $\Delta^{17}\text{O}$ as the sulphate in such depositional settings would be: (i) closer to the site of initial sulphide
771 oxidation; and (ii) subjected to less microbial sulphur cycling as is common in the marine realm (i.e.
772 gradual erasure of bulk rock $\Delta^{17}\text{O}$ anomalies by oxygen isotope exchange with water). In this regard,
773 continental sulphates hold important potential as terrestrial analogues for sulphate evaporites in
774 extraterrestrial arid environments. In such settings, the erasure/absence of $\Delta^{17}\text{O}$ anomalies predicted by
775 abiogenic processes may constitute an unexplored indicator of sulphur cycling by micro-organisms.

776 This work has shown that, in environments where there is no evidence of marine influence (and
777 sedimentological evidence against it) and where a high proportion of fluvial sulphate is thought to be
778 derived from Ca-sulphate dissolution, the age of the sulphate complex may significantly pre-date the
779 age of the deposit. In the case of the Paradox Formation and the Cedar Mesa Sandstone Formation the
780 depositional age difference may be of the order of 20 to 30 Myr, spanning the Carboniferous–Permian
781 boundary, but for other continental evaporite successions such a time gap may be longer. If the
782 interpretations herein are correct, future studies of the Paradox Formation should reveal significant
783 negative $\Delta^{17}\text{O}$ values.

784 Therefore, in terms of temporal $\Delta^{17}\text{O}$ evolution, $\Delta^{17}\text{O}$ values in the Cedar Mesa Sandstone Formation
785 may more accurately reflect the atmosphere of the late Carboniferous Period. Even fewer $\Delta^{17}\text{O}$ values
786 are known from the Carboniferous Period ($n = 13$); mid to late Pennsylvanian $\Delta^{17}\text{O}$ values range from
787 0 to -0.15‰ (Bao, 2005; Bao *et al.*, 2008). Hence, the dataset presented here provides an important new
788 constraint on late Carboniferous $\Delta^{17}\text{O}$ values.

789

790 **CONCLUSIONS**

791 By using the Cedar Mesa Sandstone Formation of the Paradox Basin as a case study, this work has
792 demonstrated that evaporites within a transitional setting between desert ergs and saline lakes exhibit
793 unique sedimentological characteristics. The relationships may be summarized by separate models of
794 erg-margin and lacustrine-margin systems. Erg-margin systems represent small-scale saline interdunes

795 hosted between and around dominantly aeolian sediments. Lacustrine-margin systems represent
796 evaporitic development in saline pan and mud flats around an arid saline lake. These settings develop
797 in response to the dominance of either an aeolian erg-margin setting or a lacustrine dominant lake-
798 margin setting due to climatic fluctuations between arid and humid conditions.

799 Sulphur isotope systematics, coupled with detailed sedimentological constraint on depositional
800 processes, shows that marine isotopic signals recorded within gypsum/anhydrite deposits are recycled
801 from older, underlying marine evaporites. This implies that the $\Delta^{17}\text{O}$ 'atmospheric fingerprint' within
802 these early Permian non-marine deposits does not date to the time of deposition. Rather, the $\Delta^{17}\text{O}$ values
803 reflect late Pennsylvanian atmospheric conditions; older $\Delta^{17}\text{O}$ values have been recycled into younger
804 evaporites. Therefore, in dryland settings, the heightened potential for sulphate weathering and
805 reprecipitation, raises the possibility that the age of a sulphate complex and the age of the host evaporite
806 deposit could be decoupled from one another. This holds important implications for studies attempting
807 to reconstruct temporal $\Delta^{17}\text{O}$ trends from non-marine evaporites.

808 There are, to the authors' knowledge, no $\Delta^{17}\text{O}$ values published from either the late Carboniferous or
809 early Permian periods. Therefore, provided the temporal caveats outlined above are applied, the $\Delta^{17}\text{O}$
810 data presented here can be used to constrain coupled biosphere-atmosphere evolution during the
811 Phanerozoic Eon. They can be included in $\Delta^{17}\text{O}$ compilations because the source of the recycled signals
812 has been constrained. This, however, may not be the case for all non-marine evaporites.

813 To avoid misleading interpretations of secular $\Delta^{17}\text{O}$ change, isotopic data must be closely integrated
814 with sedimentological processes, basin evolution, regional geology and evaporite paragenesis. Such
815 contextual work should be conducted in future $\Delta^{17}\text{O}$ studies targeting non-marine evaporites so that the
816 temporal fidelity of $\Delta^{17}\text{O}$ signals can be verified.

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826 **DATA AVAILABILITY STATEMENT**

827 The data that support the findings of this study are available from the corresponding author upon
828 reasonable request.

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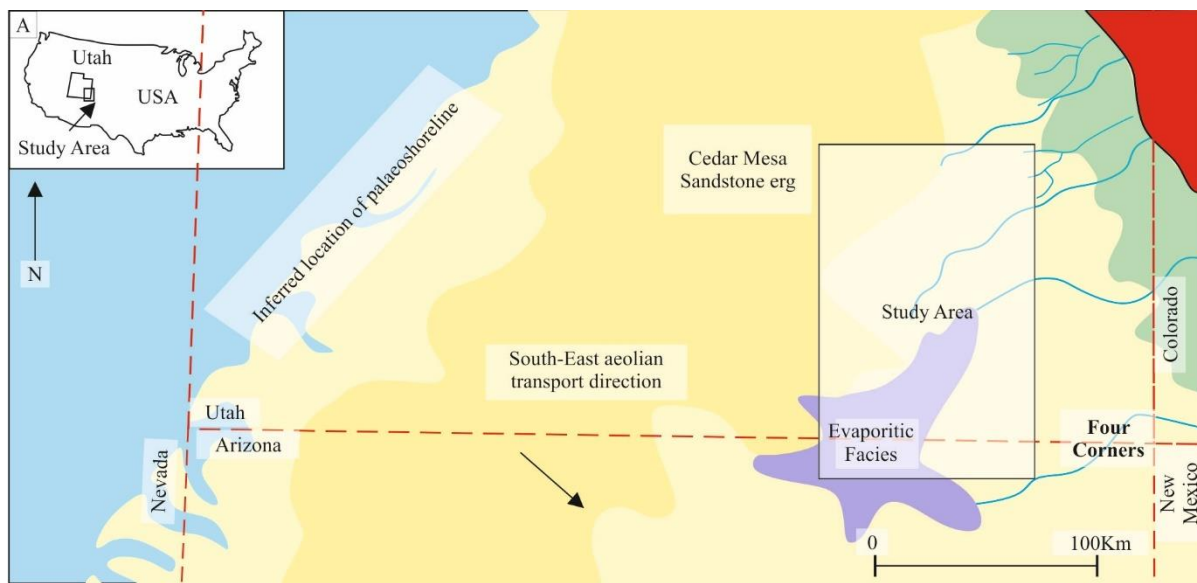
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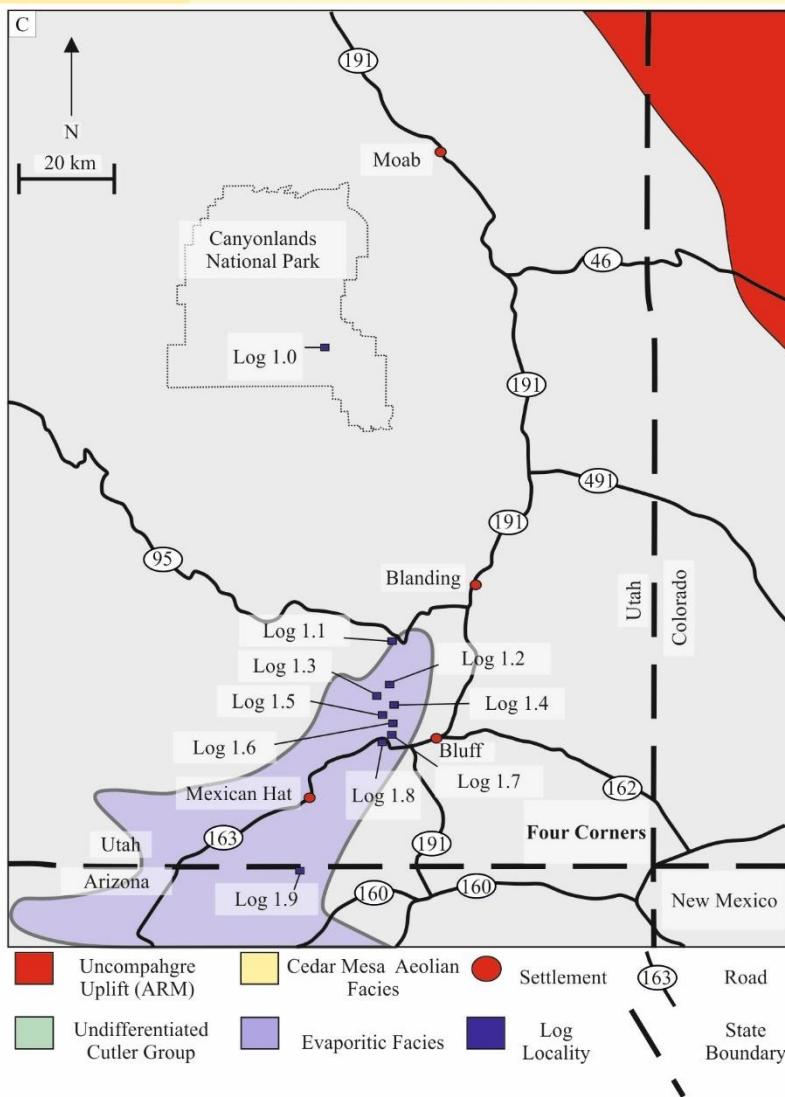
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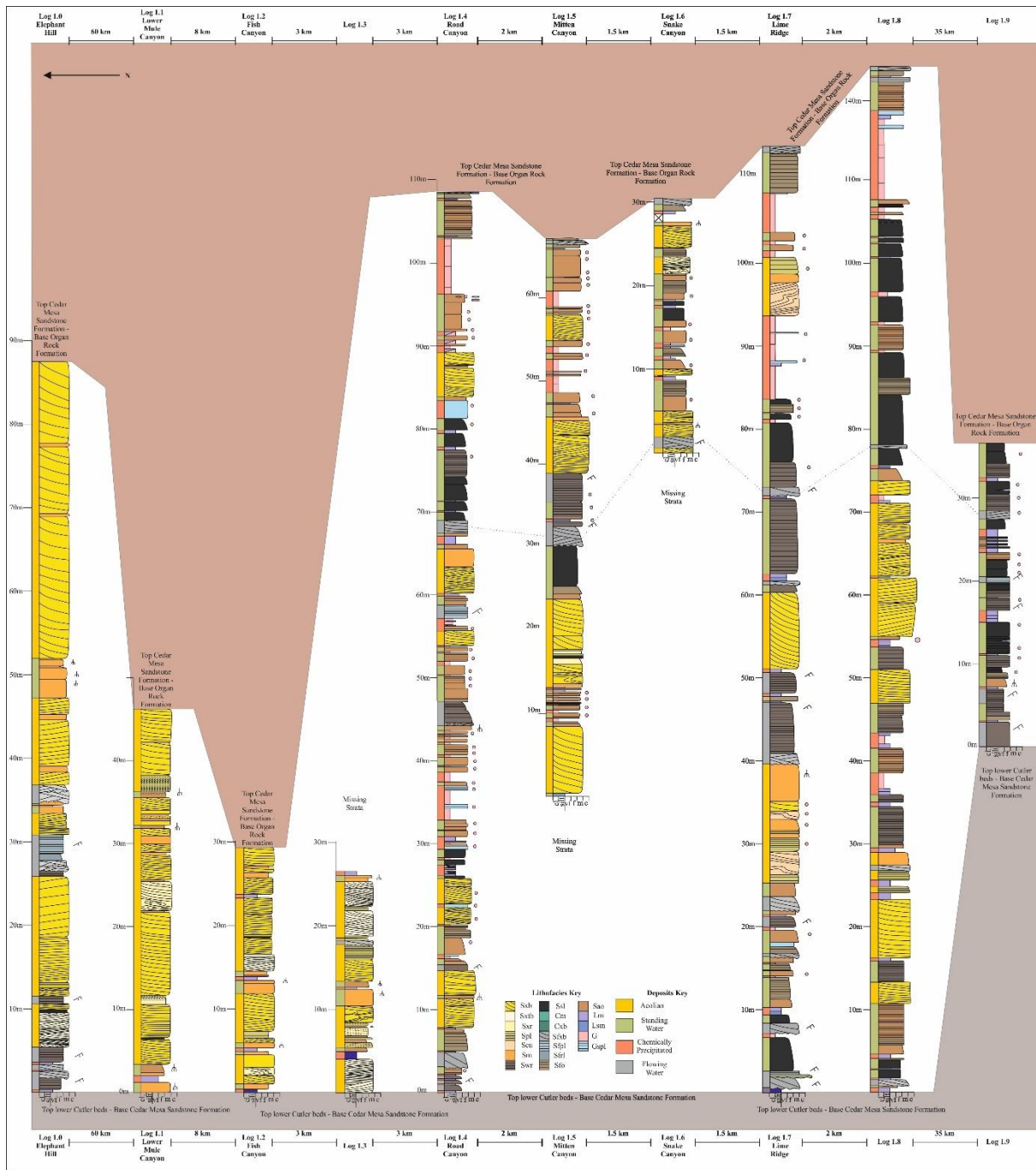
1307 **FIGURE CAPTIONS**



B	Age	Group	Formation	Depositional Setting
Triassic	251.0		Moenkopi Formation	
			De Chelly Sandstone	Aeolian dunes
Permian	299.0	Cutler Group	Organ Rock Formation	Terminal fluvial fan
			Cedar Mesa Sandstone Formation	Aeolian dunes transitioning to sabkha and lacustrine settings
			Lower Cutler Beds	Shallow marine, fluvial into loess
Pennsylvanian	309.4	Hermosa Group	Honaker Trail Formation	Shallow marine shelf and near shore environments
			Paradox Formation	Periodic deposition in restricted shallow seas. Highstand results in marine shelf sediments. Low stand resulted in sea water evaporation and precipitation of evaporites.



1309 **Figure 1 (A)** Reconstructed paleogeography of the Cedar Mesa Sandstone Formation during the early
1310 Permian Period (after Blakey *et al.*, 1988). Location of the dune field is marked in dark yellow, with
1311 the location of evaporitic sediments shown in purple against an inferred land surface (light yellow).
1312 Present day state boundaries are superimposed and the study area is highlighted. **(B)** Stratigraphy of the
1313 study area from Pennsylvanian to Triassic times. Unconformities are marked with an undulating line
1314 (after Barbeau, 2003). **(C)** Study area and log localities pictured with roads, national parks and state
1315 boundaries. Palaeogeographical location of the sabkha facies (purple) and Uncompahgre Uplift (red)
1316 are also shown (after Blakey *et al.*, 1988).



1317

1318 **Figure 2** Correlation of all logs within the study area. Logs are correlated from top of the lower Cutler
 1319 beds to base of the Organ Rock Formation. Traceable units within the Cedar Mesa Sandstone Formation
 1320 are shown with dashed lines. Logs are arranged along a roughly north – south transect. Distance between
 1321 each log is shown. Lithofacies and deposit keys are shown in the figure.

1322 **Table 1** X-ray diffraction (XRD) sample numbers against log depth, with evaporite texture and XRD
 1323 phases identified

Sample No.	Log Depth (m)	Evaporite Texture/Description	XRD phases identified
1	2	Vein within sandstone	gypsum, quartz, anhydrite
2	3	Nodule	gypsum, quartz, boracite
3	5	Vein within saline pan	gypsum, quartz, celestite
4	7.5	Vein	gypsum, quartz
5	12.5	Vein	gypsum
6	16	Interbedded with limestone	gypsum, quartz, anhydrite
7	16	Nodule within palaeosol	gypsum, quartz, anhydrite, magnesian calcite, feldspar
8	21	Vein within sandstone	gypsum, quartz
9	26.5	Interbedded gypsum and limestone	gypsum, quartz, rutile
10	27	Massive	gypsum
11	28	Massive	gypsum
12	30	Massive	gypsum, quartz
13	33.5	Enterolithic growth	gypsum, anhydrite
14	35.5	Enterolithic growth	gypsum, quartz, calcite
15	39.5	Massive	gypsum, anhydrite
16	42	Tepee structure	gypsum, quartz
17	42.5	Massive	gypsum, quartz, anhydrite
18	43	Massive	gypsum, quartz, anhydrite

19	43	Massive	gypsum, quartz
20	49	Interbedded with sandstone	gypsum, quartz, calcite
21	50	Interbedded with sandstone	gypsum, quartz, potassium sulfate
22	51.5	Interbedded with sandstone	gypsum, quartz
23	55.5	Massive	gypsum, quartz
24	65	Nodule within interdune sandstone	gypsum, quartz
25	66	Nodule within palaeosol	gypsum, quartz
26	70	Nodule	gypsum, quartz, boracite
27	70	Nodule	gypsum, quartz, anhydrite
28	89.5	Nodule within saline pan	gypsum, quartz, rutile
29	91.5	Nodular gypsum	gypsum anhydrite, muscovite
30	96.5	Massive	gypsum, celestite
31	97	Massive	gypsum
32	97.5	Massive	gypsum
33	98.5	Massive	not determined
34	99.5	Massive	gypsum
35	101	Massive	gypsum
36	102.5	Massive	gypsum

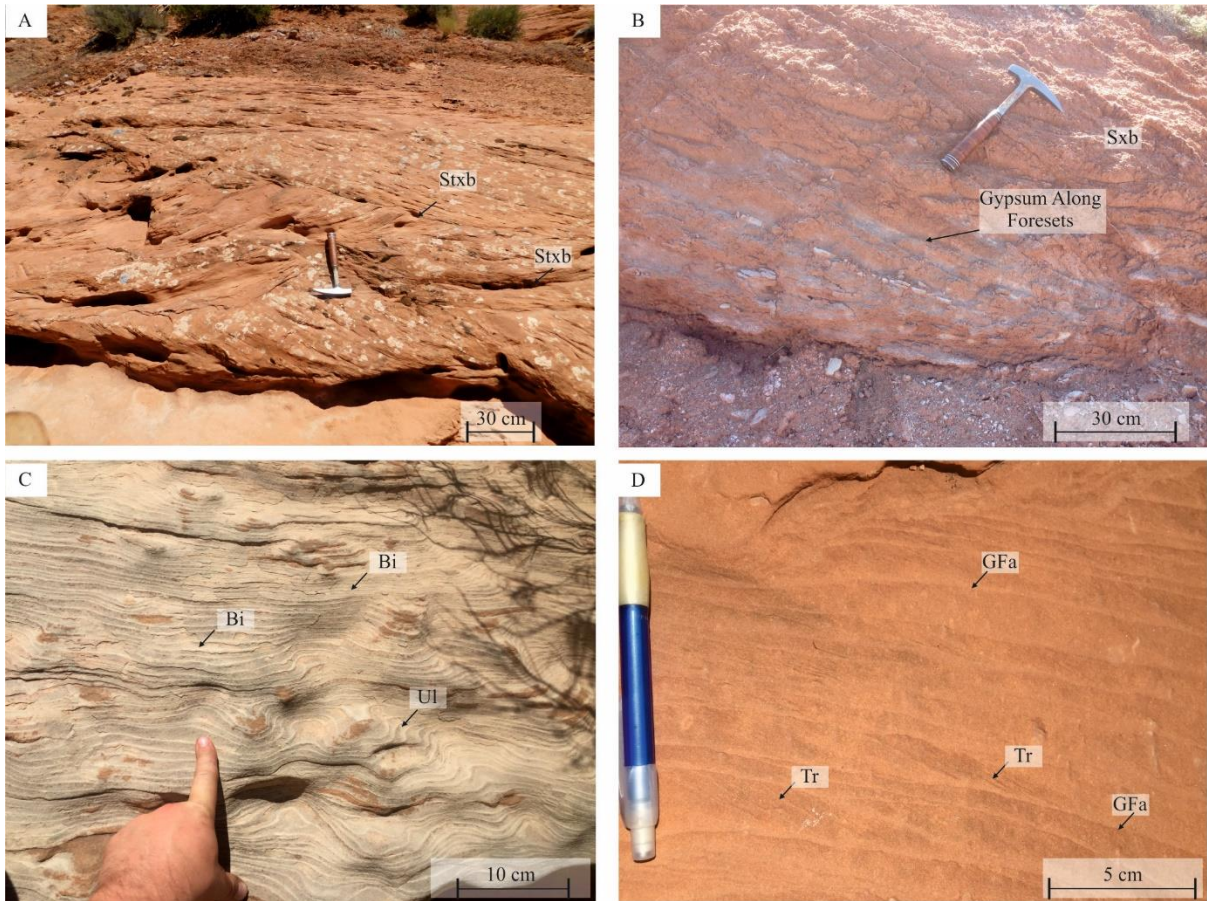
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1325 **Table 2** Lithofacies of the Cedar Mesa Sandstone Formation

Facies	Code	Lithology	Sedimentary Structures	Interpretation
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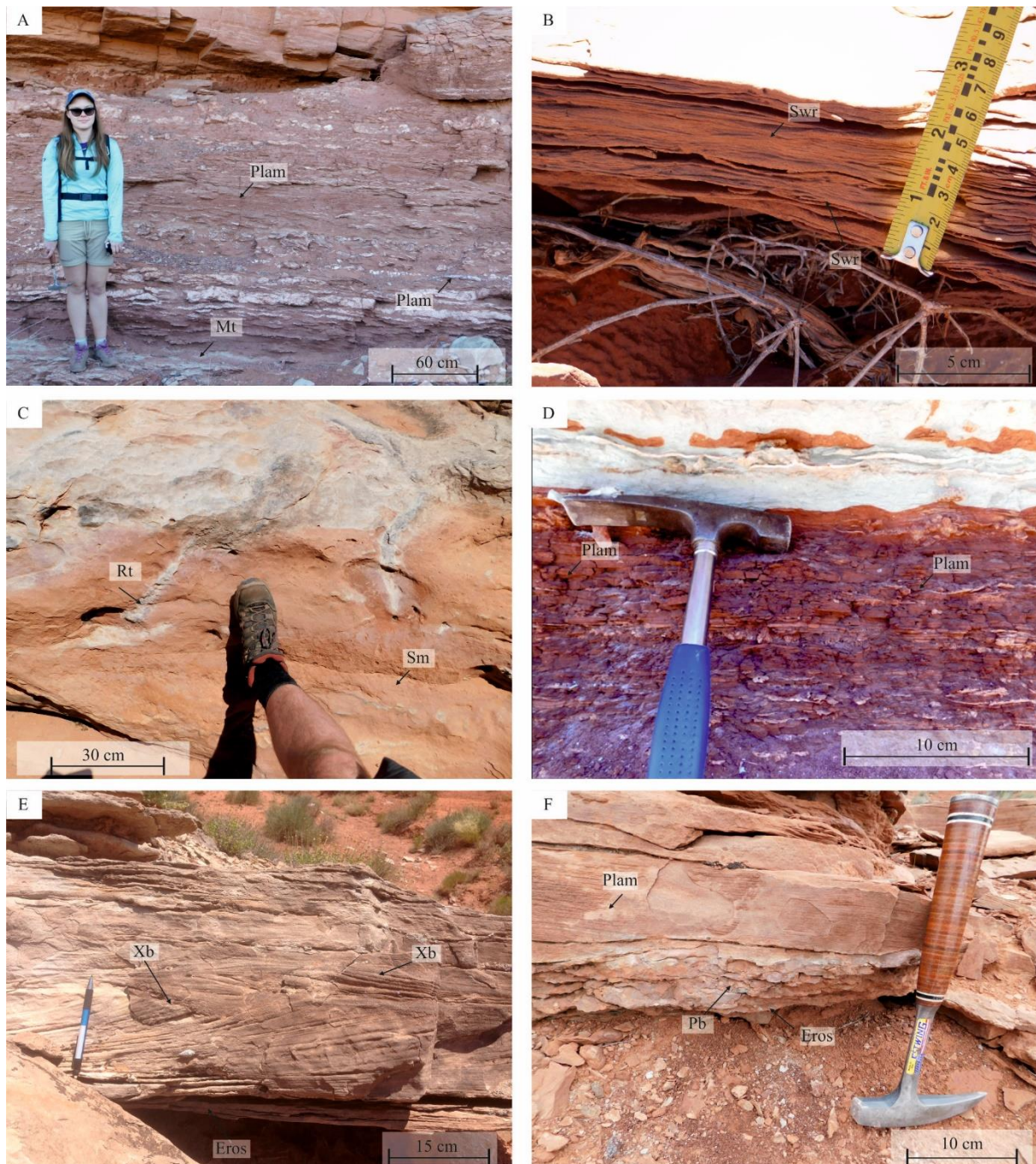
Planar cross-bedded sandstone	Sxb	Grey to orange, fine to medium-grained, well sorted & well rounded sandstone.	Planar cross-bedding with mm/cm scale alternations in grain size occurring in single or multiple sets <10 cm, with localised soft sediment deformation.	Alternating laminations of grain fall and flow and migration of wind-blown straight-crested dune-scale bedforms and dune trains. Soft sediment deformation formed as a result of slumping of dune lee slope and/or haloturbation.
Trough cross-bedded sandstone	Stxb	Grey to orange, fine to medium-grained, well sorted & well rounded, sandstone.	Trough cross-bedding with mm/cm scale alternations in grain size occurring in single or multiple sets.	Alternating laminations of grain fall and flow and migration of wind-blown sinuous-crested dune-scale bedforms and dune trains.
Ballistic ripple Sandstone	Sxr	Yellow, orange to white, fine to medium-grained, well sorted, sub to well-rounded, sandstone.	Horizontal laminations with bimodal sorting (pin stripe) with sporadic shallowly climbing (<8°) rippleform laminae.	Saltation of fine-grained sand, which accumulates along the saltation wavelength. Reptation of coarser grains over accumulated grains, results in inverse grading. Rippleform present where grain-size differentiation enables internal foreset laminae to be distinguished.
Structureless Sandstone	Sm	Grey to orange, fine to medium-grained, well sorted & well rounded, sandstone.	Structureless with localised desiccation cracks and root traces (rhizoliths).	Rapid suspension settling of wind-blown sediment in areas affected by surface water, followed by drying.
Plane parallel laminated to oscillatory current rippled siltstone and sandstone	Swr	Dark brown siltstone-to fine-grained moderately sorted & sub-rounded, siltstone and sandstone.	Parallel laminations with a sporadic undulose texture and symmetrical cross lamination.	Low energy, sub-aqueous setting, where the deposits have settled out of suspension. Undulose and oscillation-current ripples form in response to wind action on shallow waters.
Fine-grained carbonate	Lm	Dark grey to blue, carbonate mudstone/wackestone.	Sporadically blocky or stylotised with sporadic ostracod microfossils and nodular red microcrystalline chert.	Sub-aqueous precipitation of carbonate. The high matrix-content reflects a low energy in the system at the time of deposition.
Clastic-rich carbonate	Lsm	Dark grey to blue carbonate wackestone/ packstone with >10% fine grained sand within a carbonate mud matrix.	Sporadically blocky or stylotised.	Sub-aqueous precipitation of allochthonous carbonate with siliciclastic input from neighbouring environments. The lower proportion of matrix with respect to Facies Lm reflects slightly higher energy in the system at the time of deposition.

Gypsum	G	White to peach crystalline gypsum.	Massive or laminated bands of enterolithic convoluted folds or polygonal hummocks.	Precipitation from shallow saline waters and displacive growth of evaporites within saline saturated sediment.
Gypsum-bound sandstone	Gspl	Light grey, very fine to fine-grained, moderate to poorly sorted, sub-rounded sandstone within a gypsiferous matrix and cement.	Parallel laminated to massive, often contorted by small gypsum nodules.	Flow of saline fluid and subsequent precipitation of gypsum in the pore space as water evaporated at the ground surface.
Horizontally laminated mudstone and siltstone (Damp Pedogenic)	Sfo	Purple to light brown, silt to fine-grained, mudstone and siltstone.	Fining upwards horizontal laminations, mottled.	Vegetation stabilisation and sediment binding; product of palaeosol Presence of primary sedimentary features indicates a relatively immature palaeosol most likely in proximity to fluvial channels or confined bodies of water
Structureless Fines mudstone and siltstone (Dry Pedogenic)	Sfr	Light brown, grey to yellow very fine to fine-grained, moderately to well-sorted, sub-rounded mudstone and siltstone.	Structureless, sporadic gypsum and calcrete nodules and veins.	Vegetation stabilisation and sediment binding; product of palaeosol development.
Parallel laminated siltstone and sandstone	Ssl	Dark brown to black, siltstone to very fine-grained sandstone,	Structureless to faint parallel-laminations with normal grading and high organic content. sporadic mottling	Suspension fall out within low energy waters. High organic content indicates either thermal stratification or anoxic conditions.
Planar cross-bedded sandstone	Sfxb	Brown medium-grained moderately sorted & sub-rounded sandstone.	Planar cross-bedding with normal grading, in single or multiple sets, sporadic mud clasts.	Migration of straight-crested dune-scale bedforms and dune trains sub aqueously under lower flow regime conditions.
Plane parallel-stratified sandstone	Sfpl	Grey to brown, fine to medium-grained, moderately sorted, sub-rounded, sandstone	Plane parallel-stratification	Plane beds deposited under upper-flow regime conditions.
Climbing-ripple laminated sandstone	Sfrl	Grey to brown, fine to medium-grained, moderately sorted, sandstone	Sub-critically climbing asymmetrical cross-lamination	Lower flow regime ripple- scale bedform migration.



1328

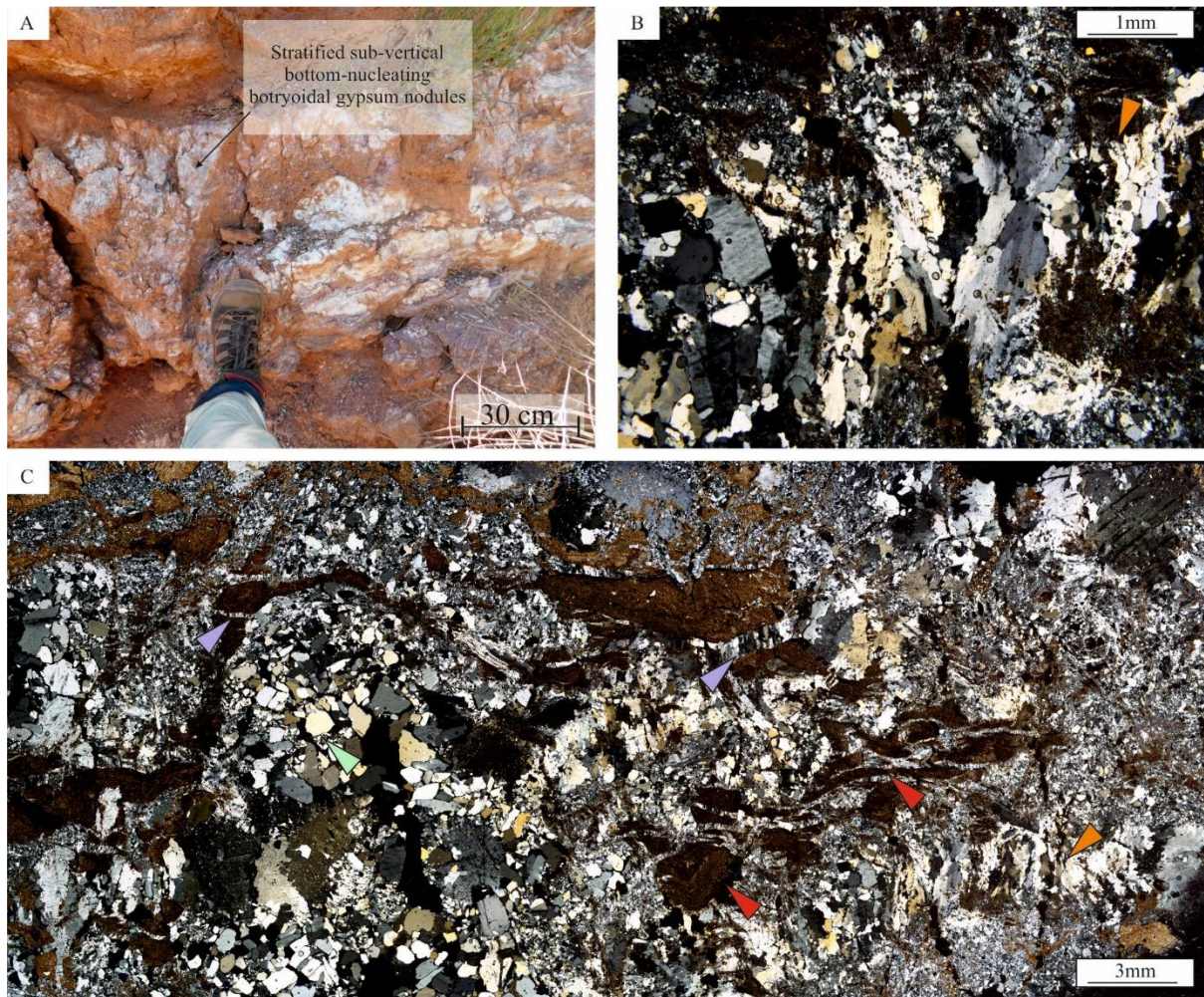
1329 **Figure 3** Key features of aeolian sediments. (A) Small trough cross bedding (Stxb) highlighted with
 1330 arrow. (B) Gypsum rich dune facies (Sxb) with white gypsum along bounding surfaces (arrowed). (C)
 1331 Ballistic ripple sandstone (Sxr) with bimodal sorting (Bi arrowed) and undulating laminations (Ui
 1332 arrowed). (D) Ballistic ripple sandstone (Sxr) with shallowly climbing rippleform laminae (Tr arrowed)
 1333 between grainfall deposits (Gfa arrowed)



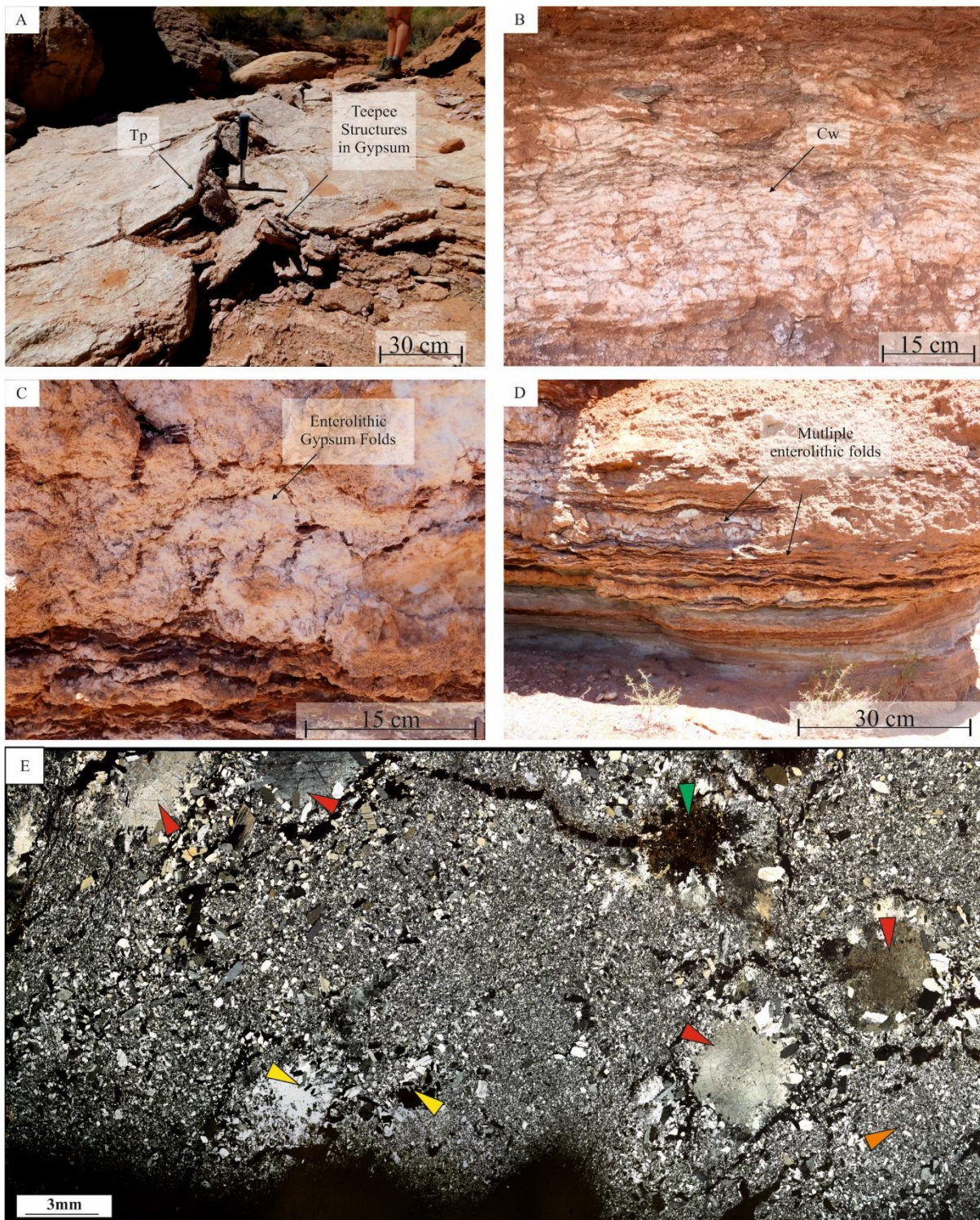
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1335 **Figure 4** The key features of standing water (A) to (D) and flowing water deposits (E) and (F). (A)
 1336 Parallel laminated siltstone and sandstone (Ssl) with planar laminations (Plam arrowed) and mottled
 1337 appearance (Mt arrowed). (B) Plane parallel laminated to oscillation-current rippled siltstone and
 1338 sandstone (Swr). (C) Structureless Sandstone (Sm) with root traces (Rt arrowed). (D) Horizontally
 1339 laminated mudstone- and siltstone (Sfo) with parallel laminations (Plam arrowed). (E) Planar cross-
 1340 bedded sandstone (Sfxb) with multiple cross-bedded sets (Xb arrowed) and erosional base (Eros). (F)

1341 Plane parallel-stratified sandstone (Sfpl) overlying thin beds of Sfxb with mud clasts (Pb arrowed) and
 1342 erosional base (Eros)



1343
 1344 **Figure 5** Key features of bedded gypsum/anhydrite deposits. (A) Bottom growing botryoidal nodules.
 1345 (B) Thin section showing undulating beds of bladed gypsum with vertical textures (orange arrow). (C)
 1346 This section showing alabastrine matrix of gypsum with: quartz sand grains (green arrow), undulating
 1347 bladed vertical gypsum (orange arrow), veins filled with satin spar (purple arrow) and brown mud clasts
 1348 (red arrow).



1349

1350 **Figure 6** Key features of displacive gypsum/anhydrite deposits. (A) Tepee structure (Tp arrow) formed
 1351 within gypsum (G). (B) Chicken wire texture in gypsum (G). (C) Enterolithic gypsum folds (arrowed)
 1352 in background gypsum (G). (D) Multiple enterolithic folds in bed of gypsum (G). (E) Thin section
 1353 showing poroplastic gypsum matrix (orange arrow) with multiple diagenetic displacive halite
 1354 crystals (red arrow) bladed gypsum cement crystals (yellow arrow) and mud clasts (green arrow).

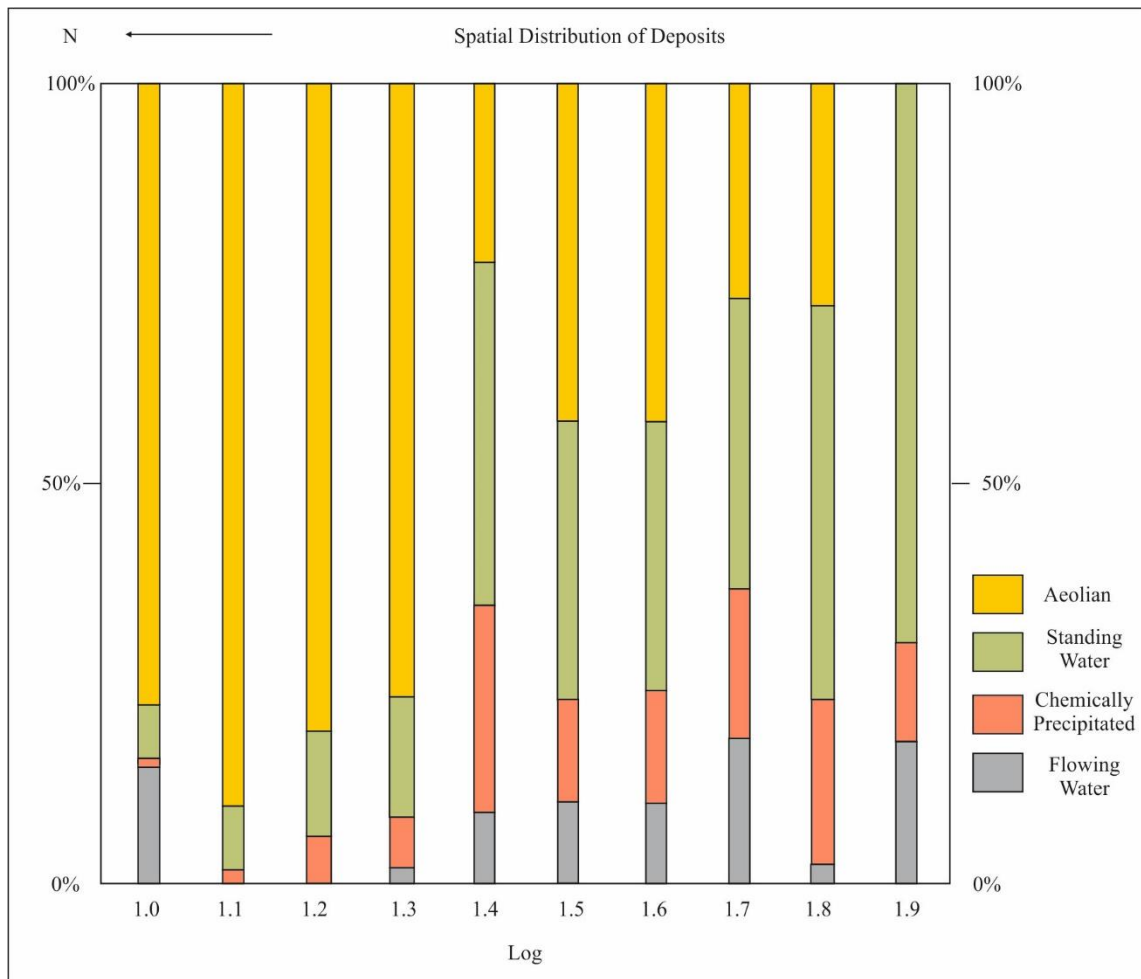


1355

1356 **Figure 7** Key features of brecciated gypsum/anhydrite deposits: (A) massive and slumped appearance

1357 in outcrop of gypsum (G); (B) crudely cross-bedded appearance of massive brecciated gypsum (G); (C)

1358 this section showing elongate gypsum crystals (orange arrow) and rounded edges (red arrow)

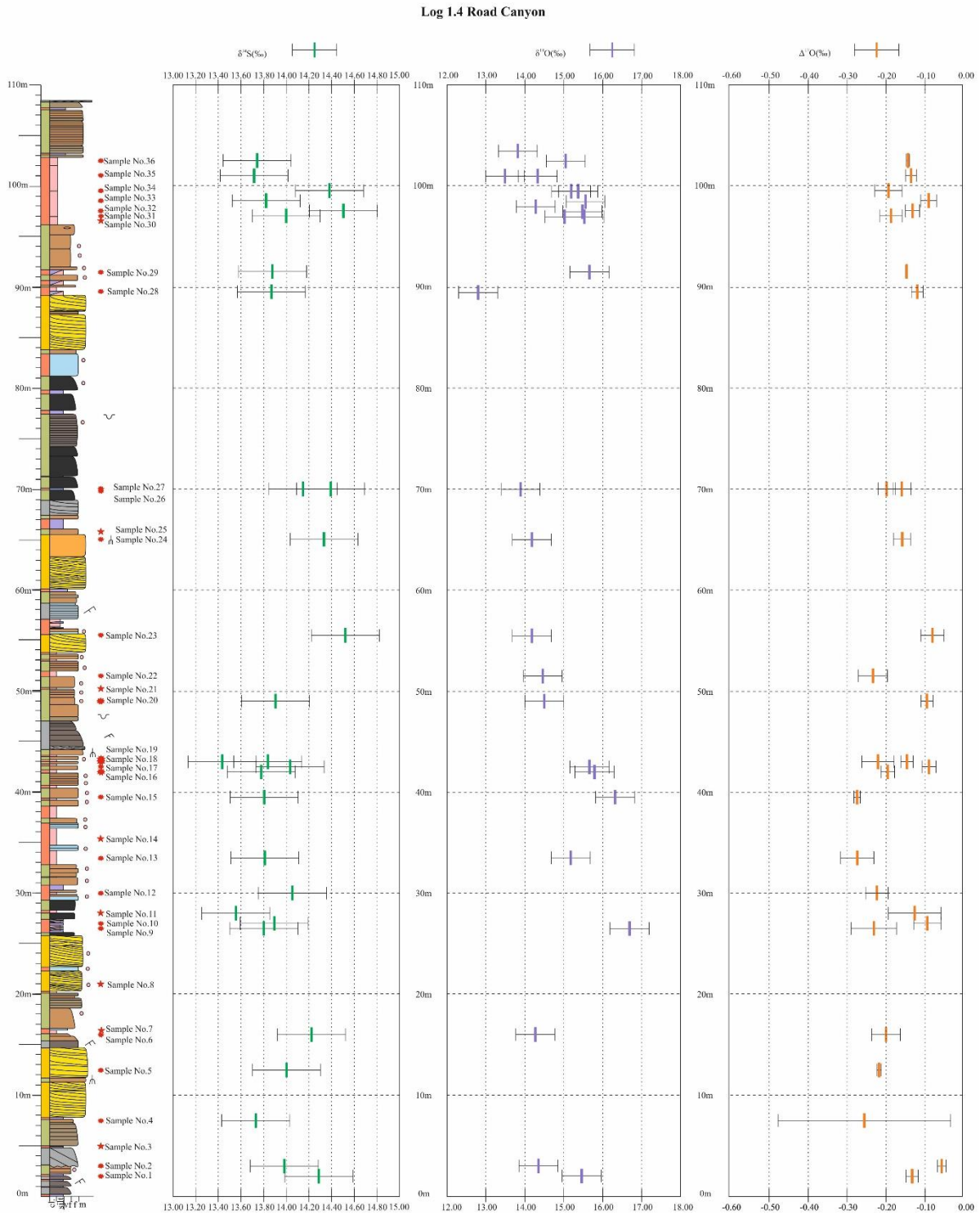


1359

1360 **Figure 8** Plot of relative proportions of aeolian, standing water, chemically precipitate and flowing
 1361 water deposits in each log across the study area. Log number is shown along the bottom (see Fig. 1 for
 1362 location) with percentage of each within each log plotted along the vertical axis.

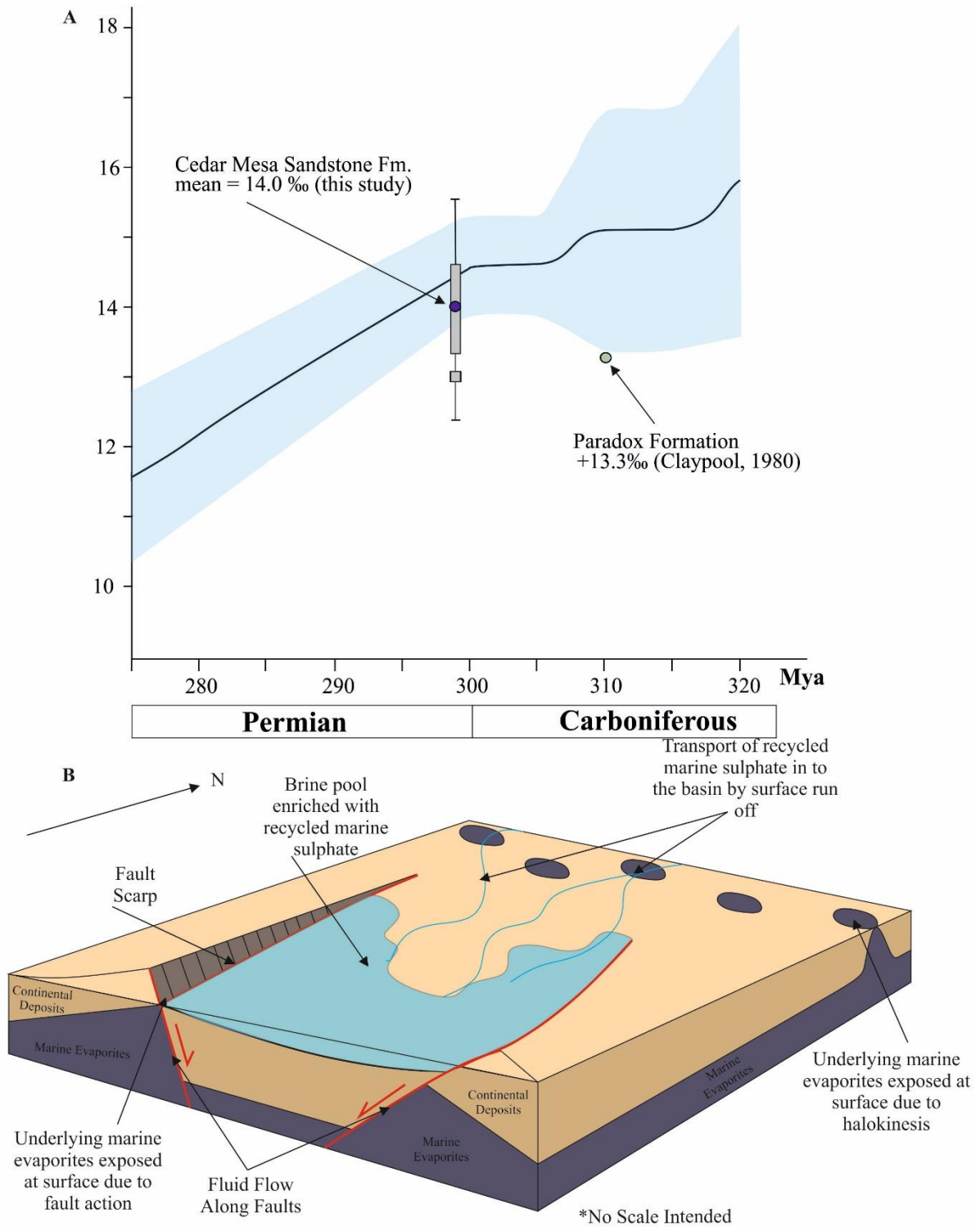
1	2	-0.13	0.02	15.46	0.50	14.29	0.30
2	3	-0.06	0.01	14.35	0.50	13.94	0.30
4	7.5	-0.26	0.22	n.d.	n.d.	13.73	0.30
5	12.5	-0.22	0.01	14.27	0.50	14.00	0.30
6	16	-0.20	0.04	16.69	0.50	14.00	0.31
9	26.5	-0.23	0.06	n.d.	n.d.	13.80	0.30
10	27	-0.09	0.03	15.18	0.50	13.88	0.30
11	28	-0.13	0.07	16.32	0.50	13.55	0.30
12	30	-0.22	0.03	15.79	0.50	13.87	0.30
13	33.5	-0.27	0.04	15.66	0.50	14.62	1.14
15	39.5	-0.27	0.01	n.d.	n.d.	13.93	0.30
16	42	-0.20	0.02	14.50	0.50	14.03	0.36
17	42.5	-0.09	0.02	14.46	0.50	13.92	0.30
18	43	-0.22	0.04	14.18	0.50	13.57	0.38
19	43	-0.15	0.02	14.18	0.50	13.43	0.30
20	49	-0.10	0.02	13.89	0.50	13.90	0.30
22	51.5	-0.23	0.04	12.80	0.50	n.d.	n.d.
23	55.5	-0.08	0.03	15.66	0.50	14.53	0.30
24	65	-0.16	0.02	15.02	0.50	14.33	0.30

26	70	-0.20	0.02	15.56	0.50	14.22	0.30
27	70	-0.16	0.02	15.19	0.50	14.39	0.30
28	89.5	-0.12	0.02	14.33	0.50	13.89	0.30
29	91.5	-0.15	0.00	15.05	0.50	13.88	0.30
31	97	-0.19	0.03	15.53	0.50	14.02	0.30
32	97.5	-0.13	0.02	15.48	0.50	14.50	0.30
33	98.5	-0.09	0.02	14.28	0.50	13.82	0.30
34	99.5	-0.19	0.04	15.37	0.50	14.38	0.30
35	101	-0.14	0.01	13.49	0.50	13.72	0.30
36	102.5	-0.14	0.00	13.82	0.50	13.74	0.30



1368

1369 **Figure 10** Isotope results plotted against log 1.4. $\delta^{34}\text{S}$ is plotted in green with error bars, $\delta^{18}\text{O}$ is plotted
 1370 in purple with error bars. $\Delta^{17}\text{O}$ is plotted in orange with error bars. Red circle next to sedimentary log
 1371 indicates sample location of isotopic and XRD data point, star indicates XRD sample only. For
 1372 sedimentary facies key see Fig. 2.



1373

1374 **Figure 11 (A)** Mean $\delta^{34}\text{S}$ of the Cedar Mesa Sandstone and Paradox Formation plotted against marine
 1375 range during the Carboniferous and Permian (blue) (after Kampschulte & Strauss, 2004). **(B)** Schematic
 1376 diagram showing possible pathways for marine sulphate flux into the continental basin via fault action,
 1377 or through halokinesis and surface run off.

